

Atmospheric Monitoring with LIDAR Systems at the Pierre Auger Cosmic Ray Observatory

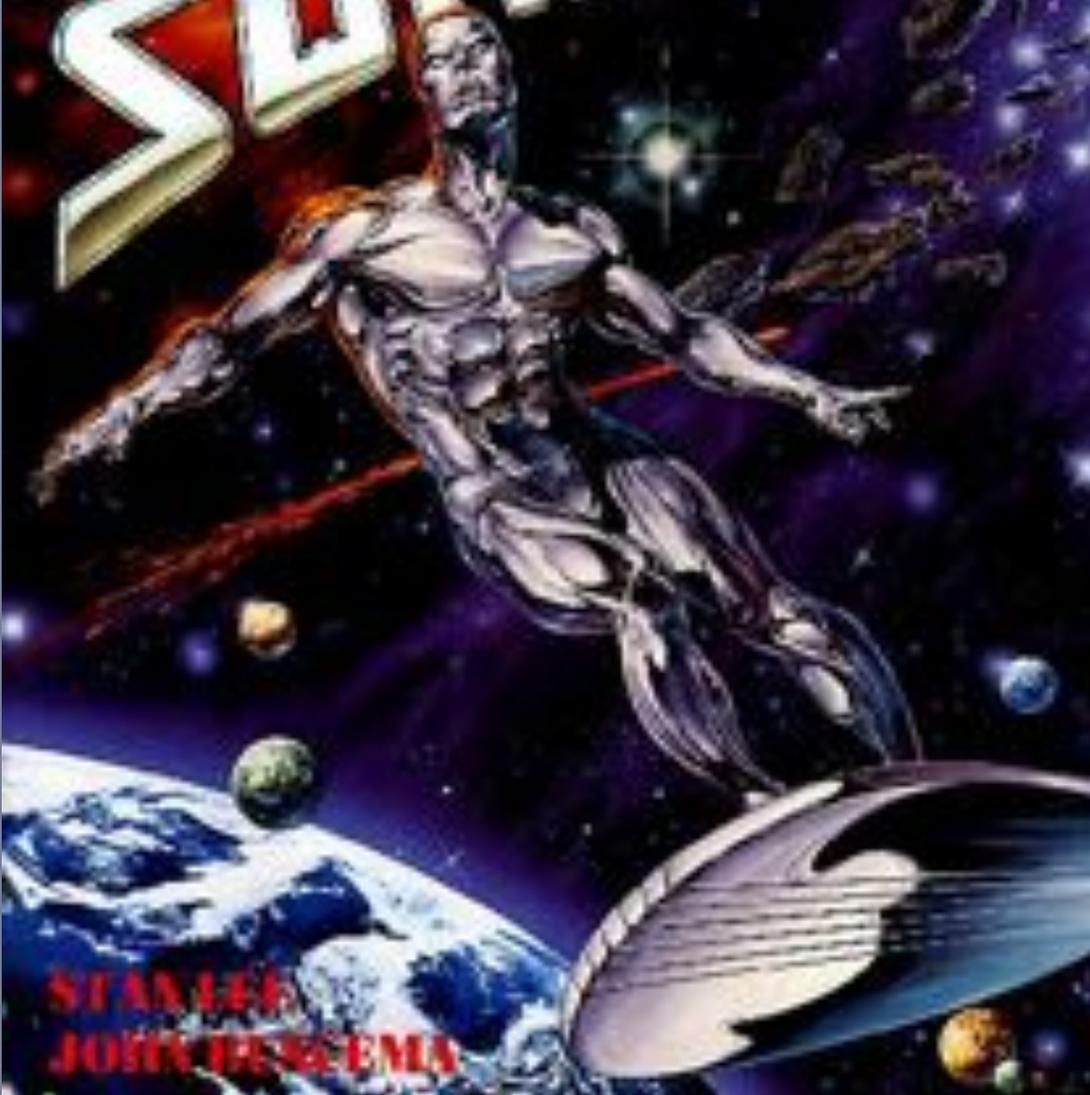
Prof. Gregory Snow



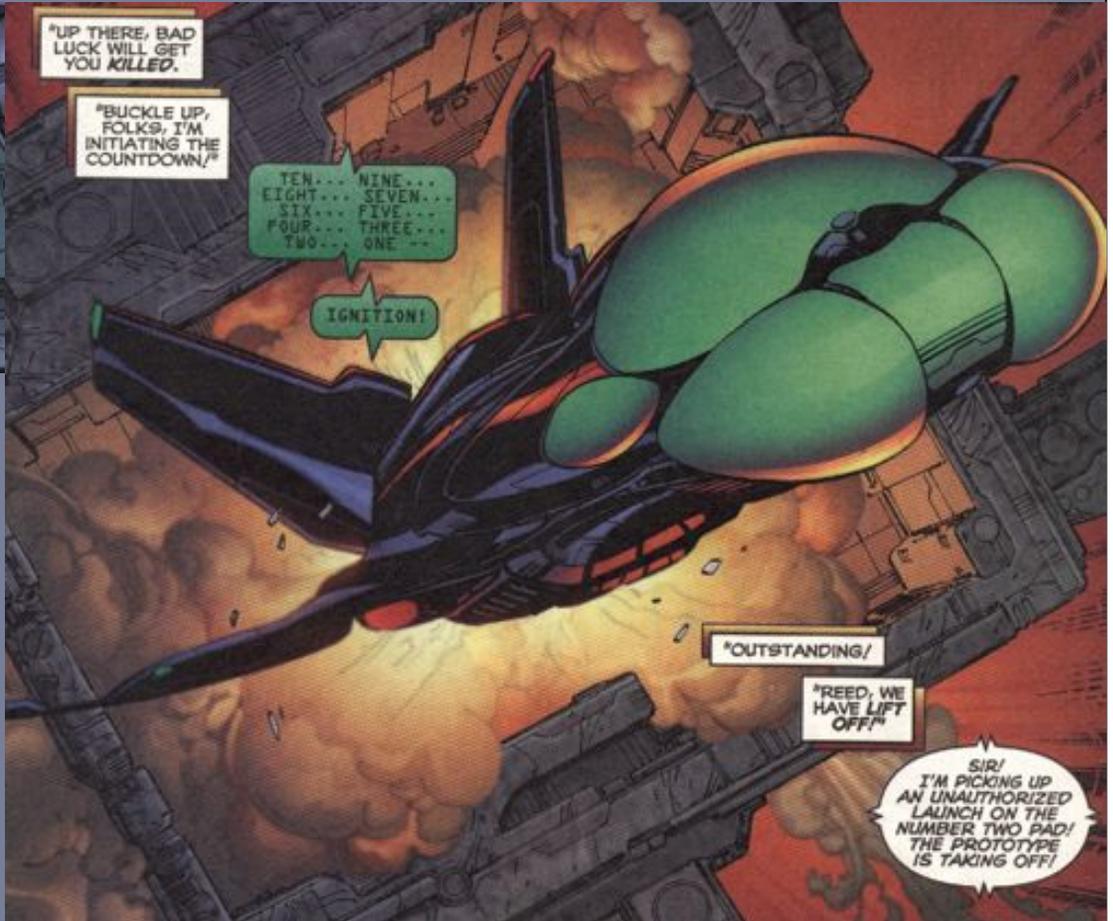
NASA meeting, Omaha: 9/17/2010

MARVEL GRAPHIC NOVEL

SILVER SURFER



STAN LEE
JOHN DE SEENA



The Fantastic Four ®
©1996 Marvel Comics

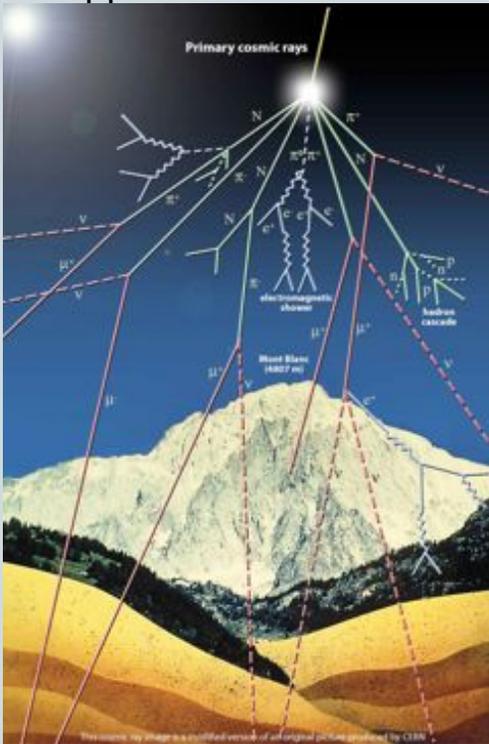


Outline

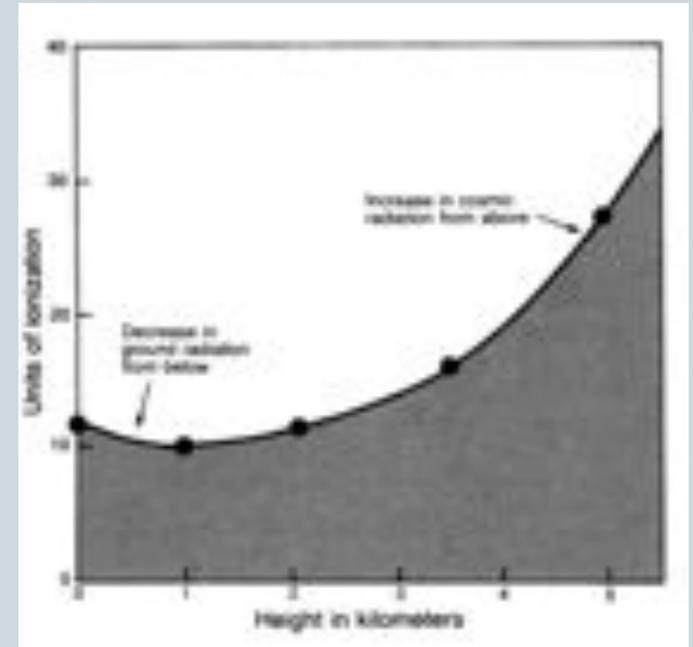
- Cosmic Ray History
 - Discoveries of Cosmic Rays & Extensive Air Showers
- Pierre Auger Cosmic Ray Observatory
 - The experiment & location
 - Surface Detectors & Fluorescence Detectors
- LIDAR Systems
 - Operation and design
- Near-Field Upgrade to LIDAR Systems by UNL Group
 - Detector Designs and Installation
 - Performance Evaluation
- Repaired Laser Study
 - Experimental Procedures
- Enhancements to online Introductory Physics courses using NASA education materials

Cosmic Ray History

- **Victor Hess** discovered cosmic rays in 1912.
 - “Rays of very great penetrating power are entering our atmosphere from above.” -



Cosmic rays are protons or large nuclei.



- In 1938 **Pierre Auger** discovered extensive air showers.
 - Some of the primary cosmic ray energy is converted into mass in the creation of new particles according to $E=mc^2$.

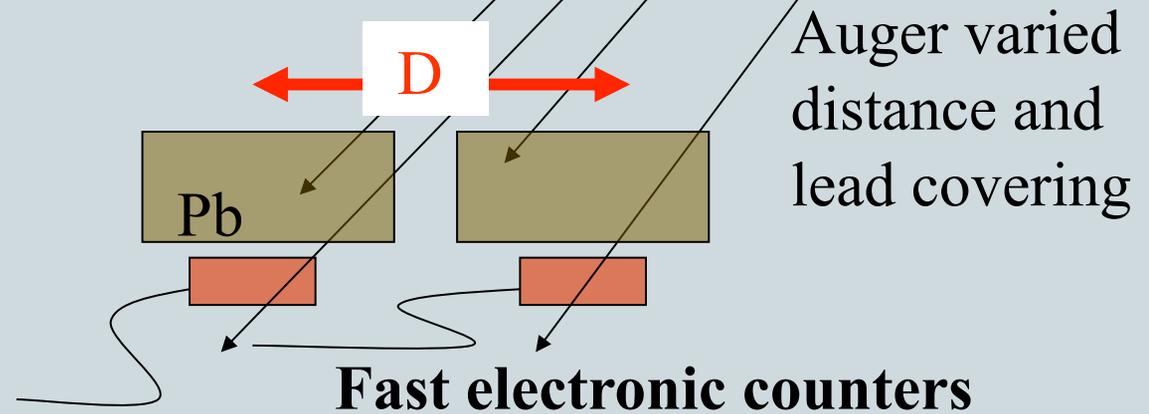
Historical high points

1912
Victor Hess
discovers
“penetrating radiation”
from space



Nobel prize 1936

1938
Pierre Auger discovers
extensive air showers





Snowmass 2001

**Replica of
Hess' Electroscope**

**Portable Geiger
Counters**



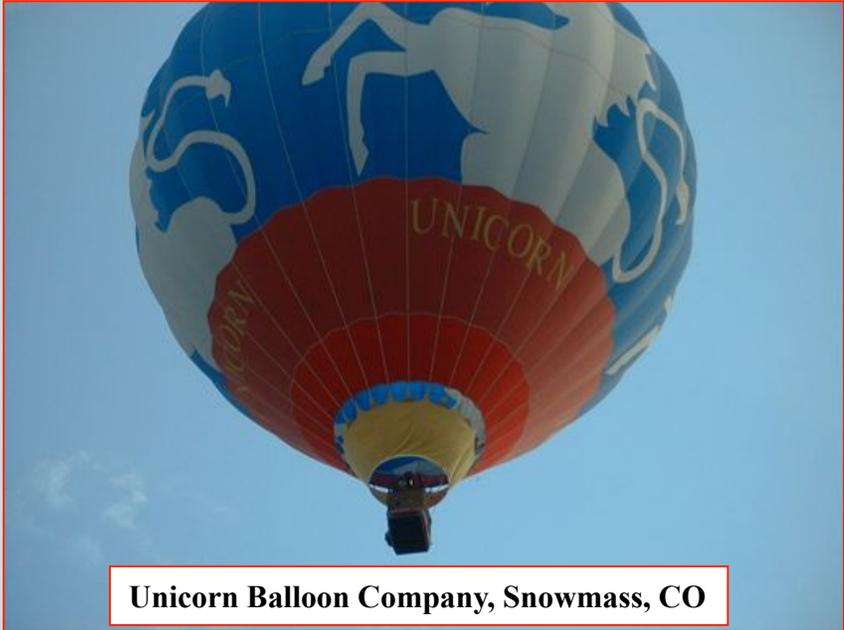
**Wilkes in
Hessian
Outfit**





**Crowd gathers
to watch
Victor Hess
flight reenactment**

Lift off !

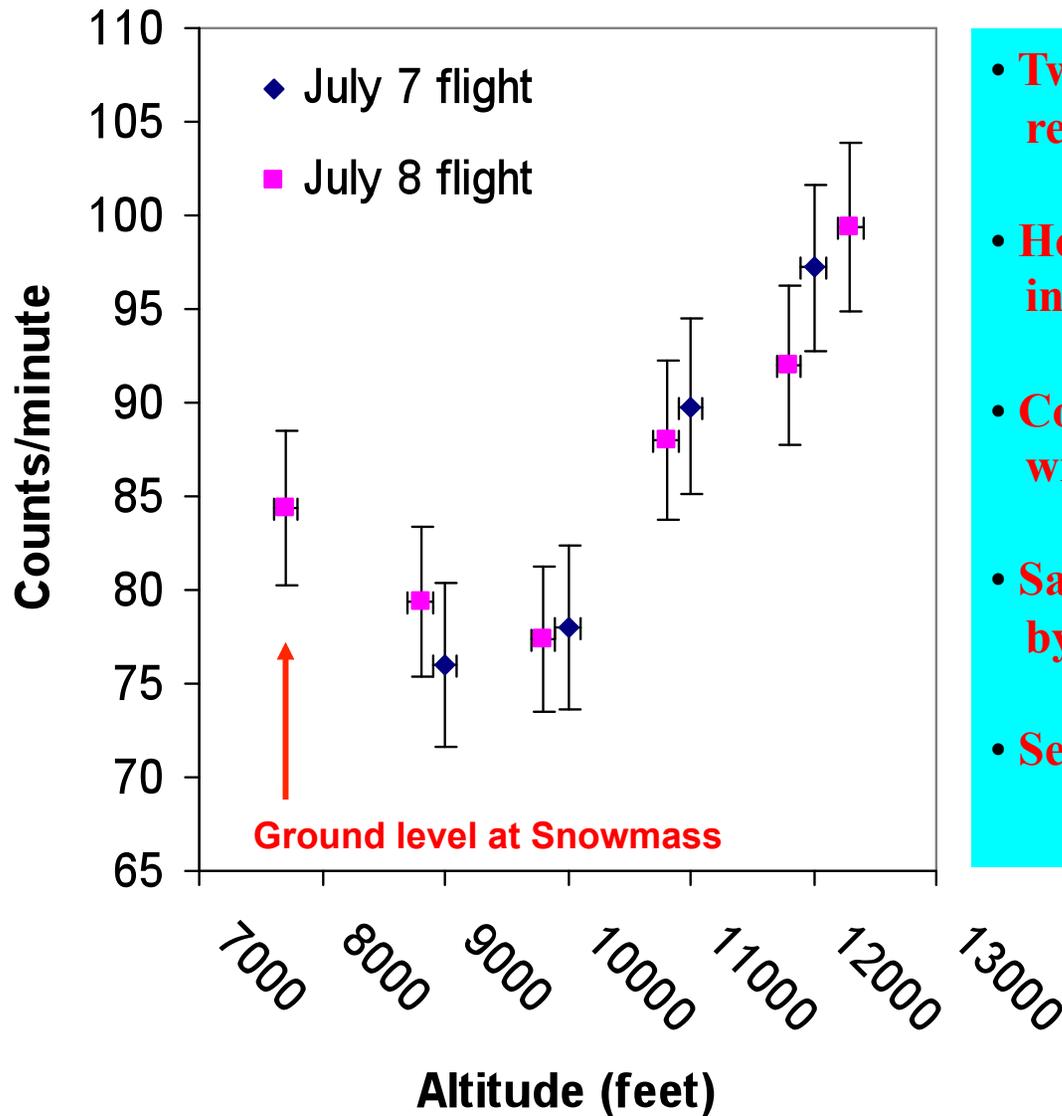


Unicorn Balloon Company, Snowmass, CO



**Data transmitted live
to ground via radio**

Snowmass Balloon Flight 2001



- **Two flights with consistent results**
- **Hovered at 1000 ft increments in altitude for 5 minutes**
- **Cosmic ray rates measured with portable Geiger counters**
- **Same effects observed by Victor Hess**
- **See FermiNews, July 27, 2001**

**Much bigger effect:
Northwest airlines
at 35,000 feet !**

The Pierre Auger Observatory



**Auger
Observatory
is here.**

Malargue is a small town on the high plains not far from a ski area in the Andes.



The Auger Collaboration

67 Institutions, 369 Collaborators

Argentina	Netherlands
Australia	Poland
Bolivia*	Portugal
Brazil	Slovenia
Czech Republic	Spain
France	United Kingdom
Germany	USA
Italy	Vietnam*
Mexico	

True International Partnership - *by non-binding agreement -*

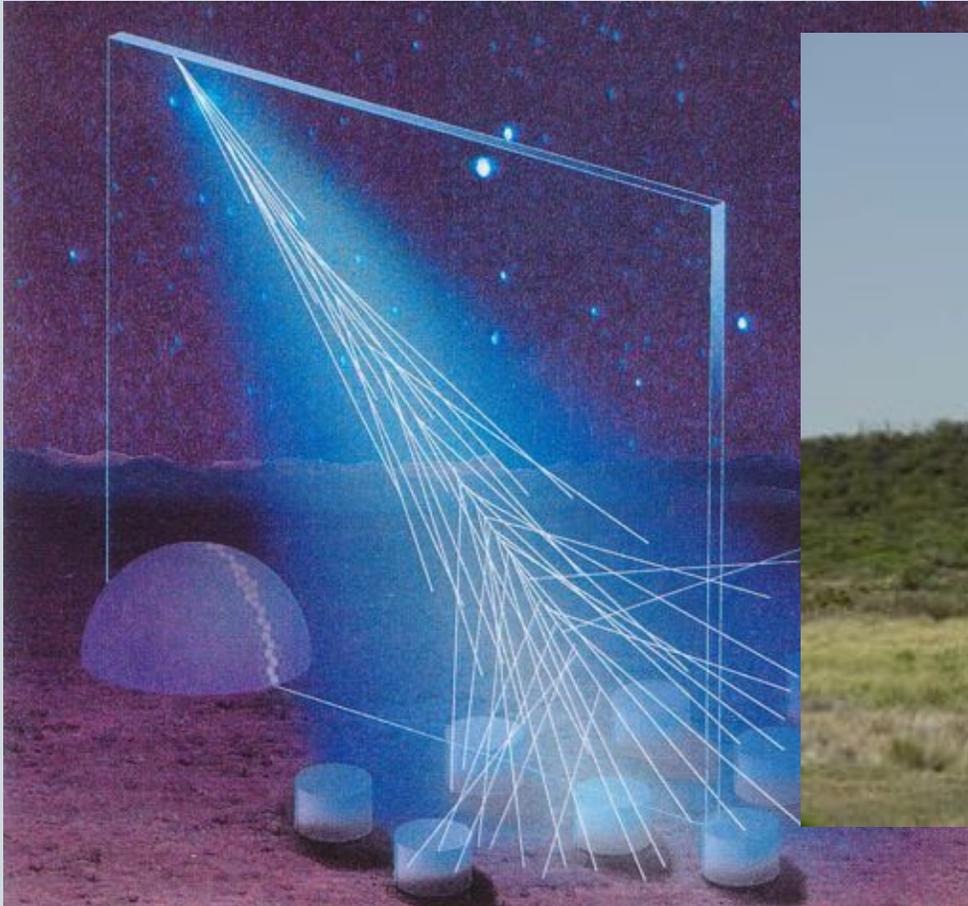
No country, region or institution dominates – No country contributes more than 25% to the construction.

* *associate*

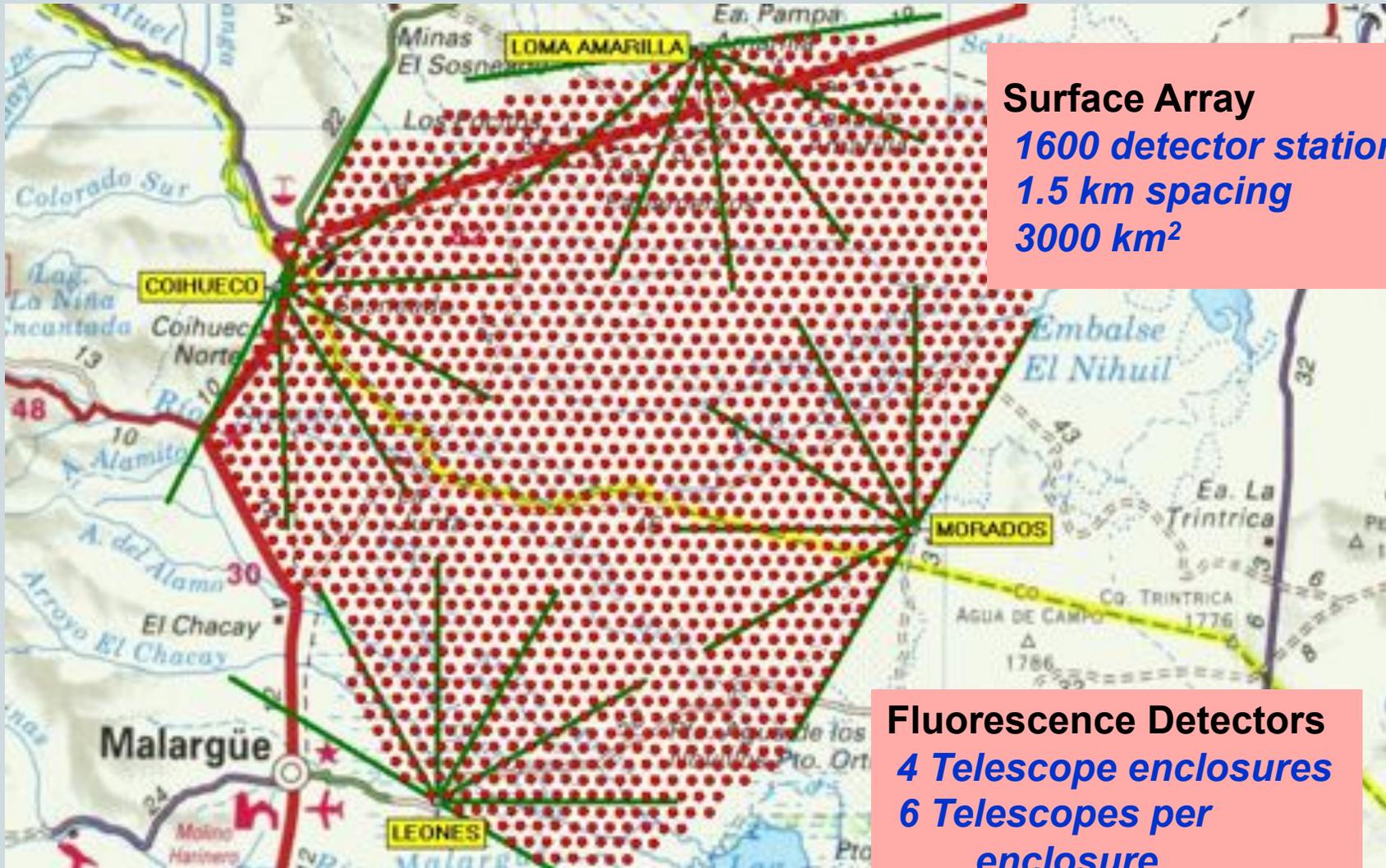


The Hybrid Design

Surface detector array + Air fluorescence detectors
A unique and powerful design



The Observatory Layout



Surface Array
1600 detector stations
1.5 km spacing
3000 km²

Fluorescence Detectors
4 Telescope enclosures
6 Telescopes per enclosure
24 Telescopes total

Surface Detectors

Detector Station



Depl



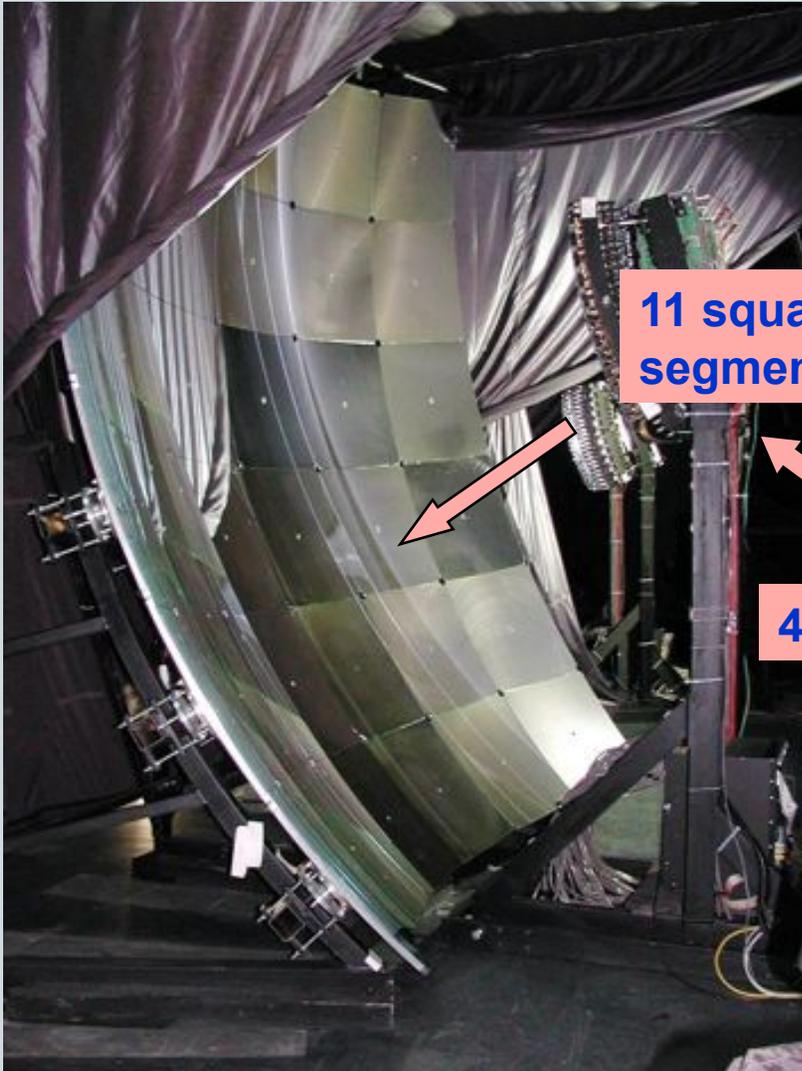
Aerial Photos of Fluorescence Buildings November 2006



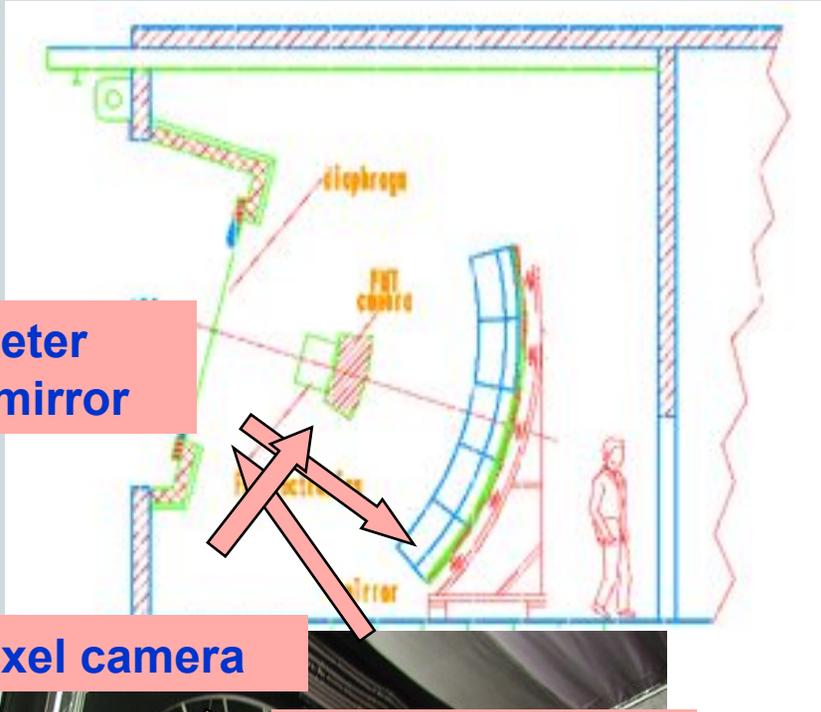
LIDAR
Enclosure



Fluorescence Detectors

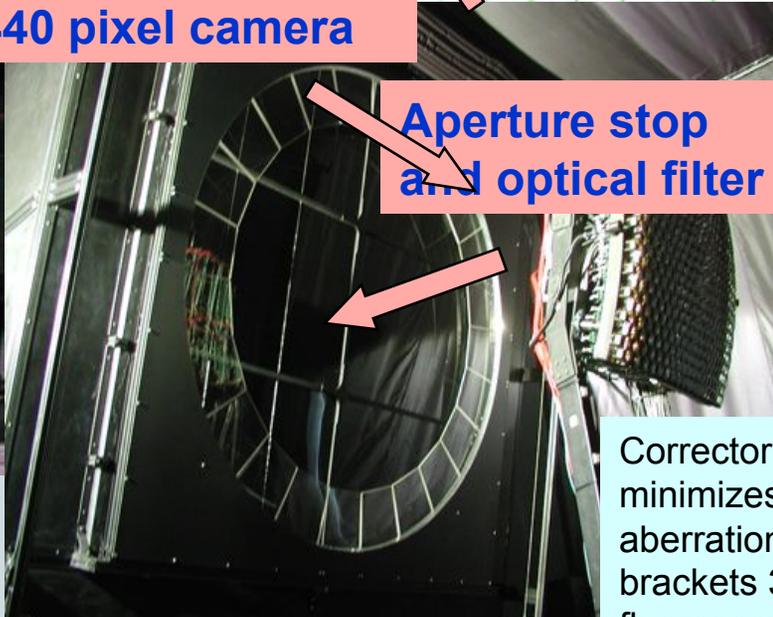


11 square meter segmented mirror



440 pixel camera

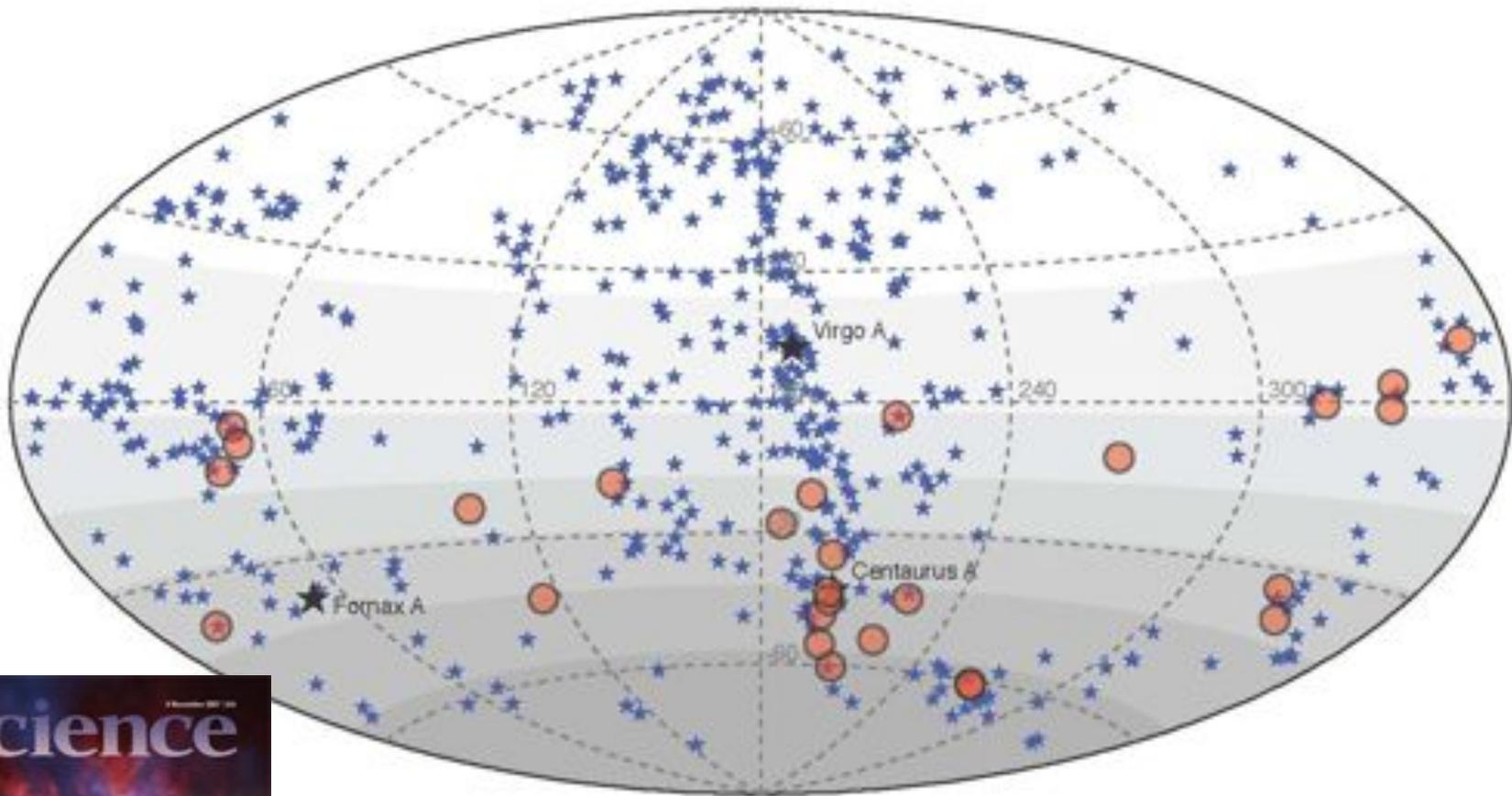
Aperture stop and optical filter



Corrector lens minimizes spherical aberrations, filter brackets 350 nm fluorescence light

FD telescopes in closed environment

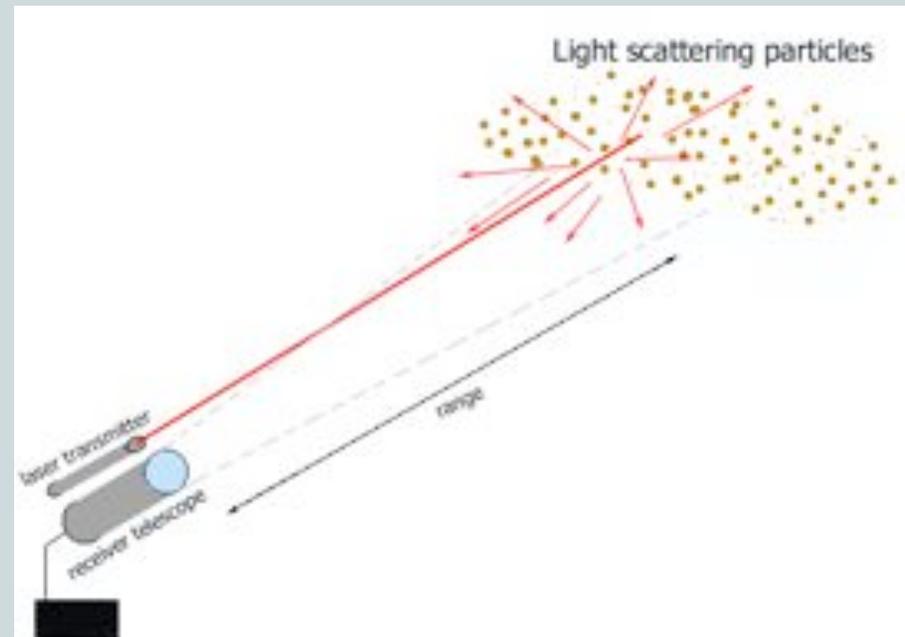
Discovery!



Science, November 2007
Arrival directions of the highest energy cosmic rays
point back to certain galaxies with “Active Galactic Nuclei”

Atmospheric Monitoring with LIDAR Systems

- Light Detection And Ranging: “LIDAR”
- LIDAR systems operate as follows:
 - UV laser is triggered to shoot pulses into the atmosphere.
 - Aerosols, molecules, and clouds in the air cause some of the laser light to be backscattered towards the LIDAR station.
 - Spherical mirrors collect backscattered laser light
 - PMT’s turn the signal into an electric current signal which is recorded as a function of time.
- Analysis of data allows for the development of atmospheric density profiles.



LIDAR enclosure near Fluorescence Detector building



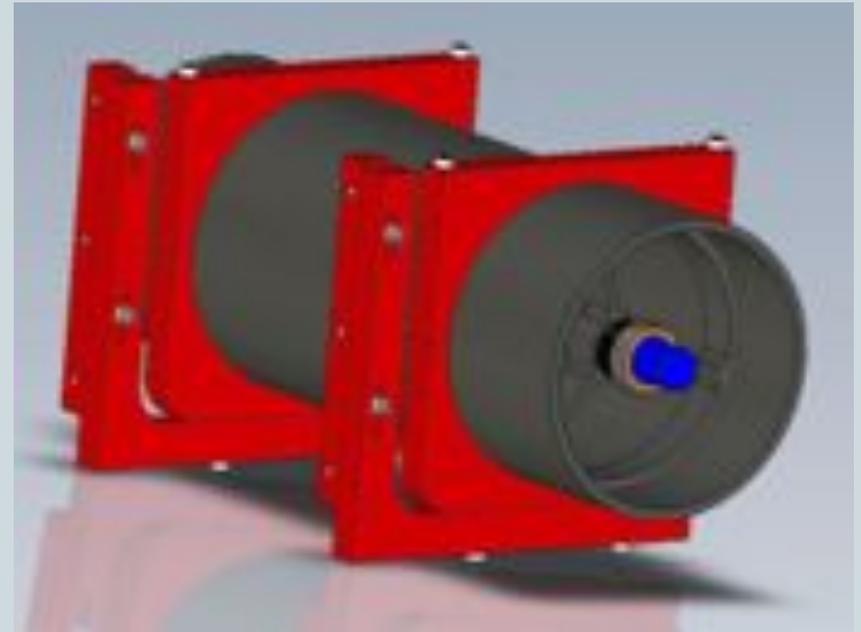
LIDAR Systems at Pierre Auger

- Original LIDAR system design had one laser box surrounded by three mirrors, each with a diameter of 80 cm.
- The mirrors and laser box are mounted to a rigid frame, which is steered and rotated in all directions to scan the entire sky above the observatory.
- After analyzing preliminary data it was determined that **upgrades needed to be made in order to improve atmospheric measurements in the near-field, that is, less than 1 kilometer in front of the laser.**
- UNL HEP Pierre Auger group:
 - Dr. Greg Snow (faculty)
 - Emily Petermann (graduate student)
 - Maria Becker (undergraduate)



Near-Field Prototype LIDAR Design

- Weak signal strength in the near-field is due to the large offset between the axis of the laser and the axes of the mirrors
 - 170 cm and 120 cm offsets
- To reduce offset the near-field detectors were designed with smaller-diameter mirrors: 15 cm diameter, focal point: 30 cm
- Simulated several detector designs with a ray-tracing program
- Optimal design was modeled in Solid Works (a 3-D design software package) and constructed by staff in the UNL Dept. of Physics Instrument Shop
- The prototype detector was installed at the Coihueco LIDAR station in Spring 2009

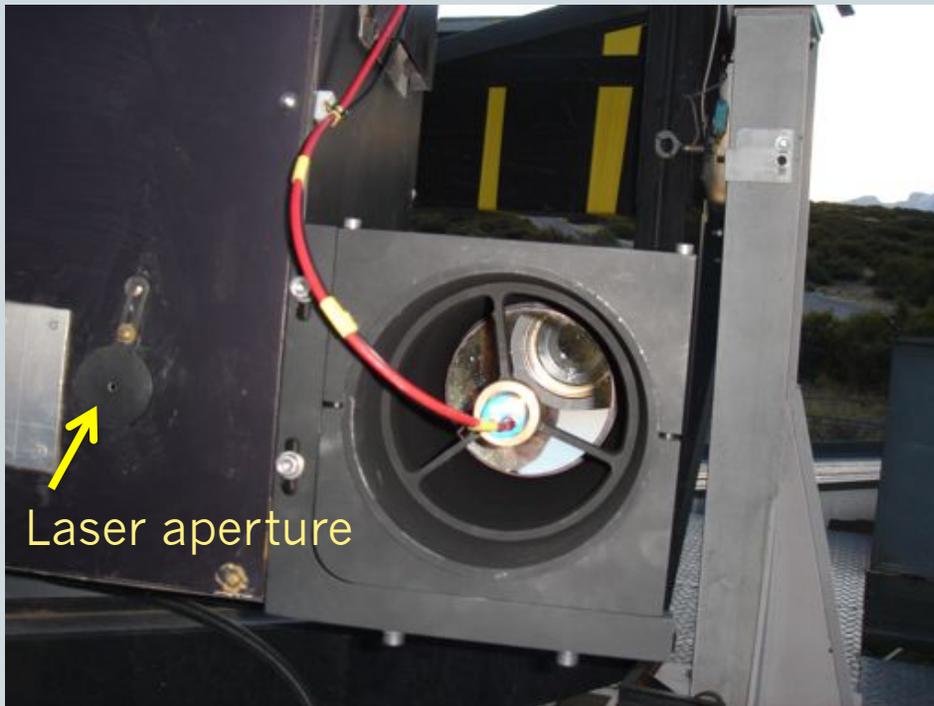


- Aluminum tube houses mirror and PMT
- Mounting support brackets allow for transverse horizontal and vertical adjustment
- Push-pull mechanism for adjusting mirror position

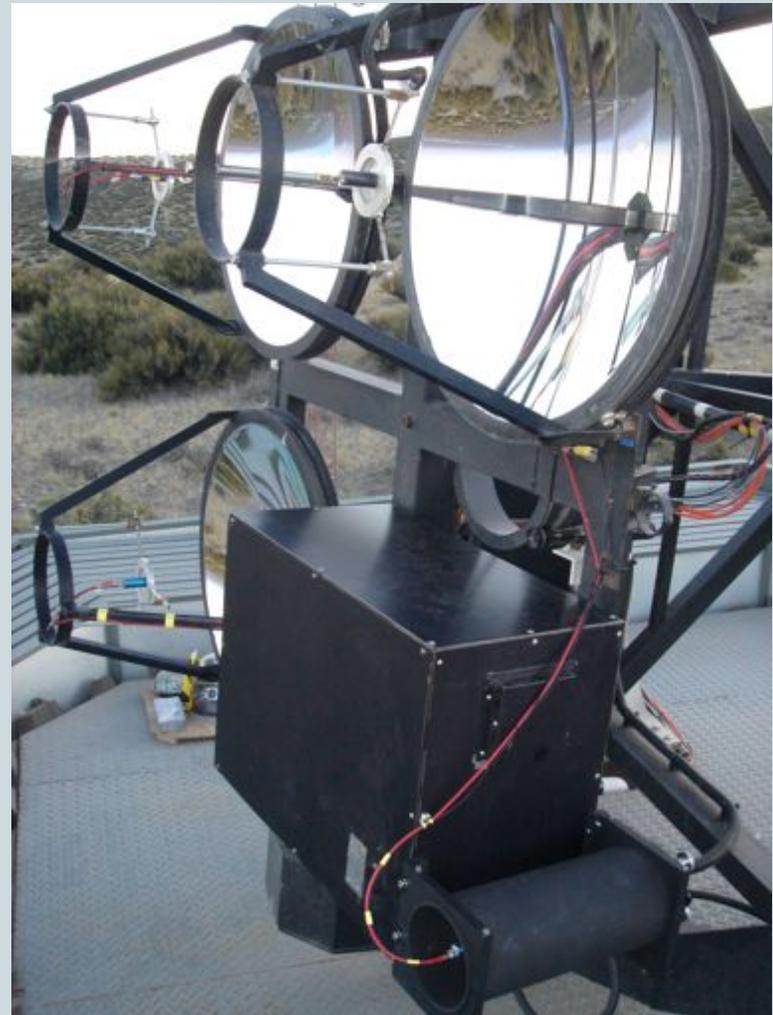
LIDAR enhancement prototype at UNL Physics instrument shop



Near-Field Prototype LIDAR Detector Installation at Coihueco

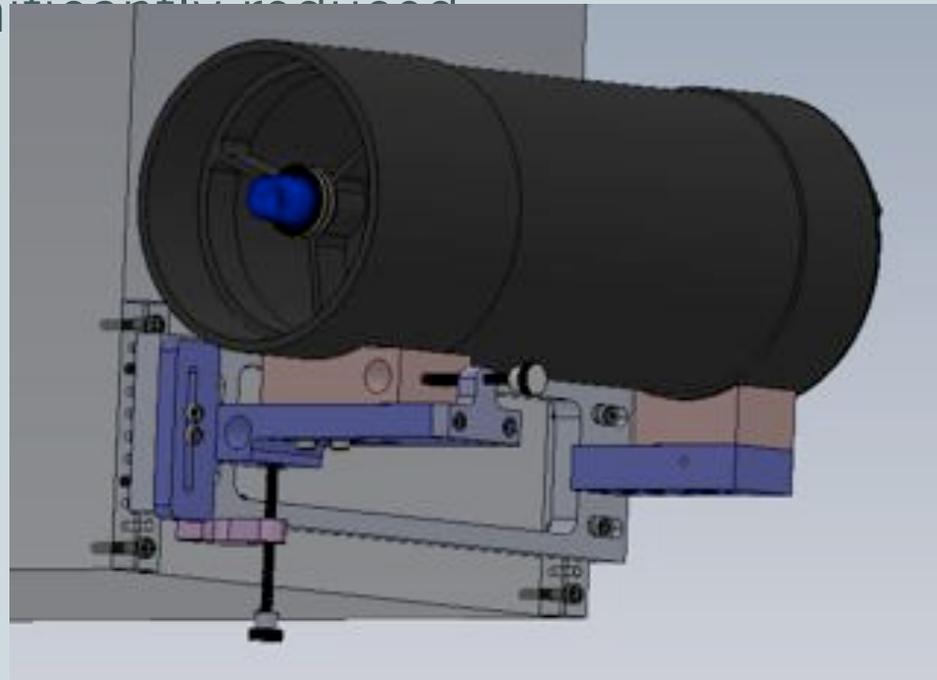


- Offset between laser aperture and mirror axis reduced to ~20 cm



Modified Near-Field Detector Design

- Prototype detector successfully collects atmospheric measurements as close as 200 m from the detector. **However, the alignment process was difficult and it was very heavy.**
- Re-designed detector's mounting and positioning structures to have with two rotational degrees of freedom.
- The weight was also significantly reduced.
- Internal detector components are the same as the for the prototype design.
- One modified near-field detector was completed in November 2009 and installed at the Loma Amarilla LIDAR station in March 2010.





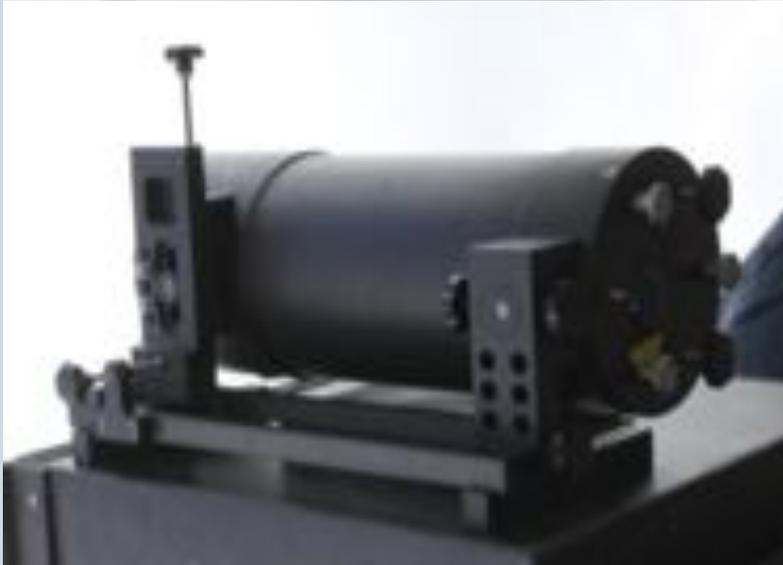
Improved LIDAR Detector



Installation in Argentina



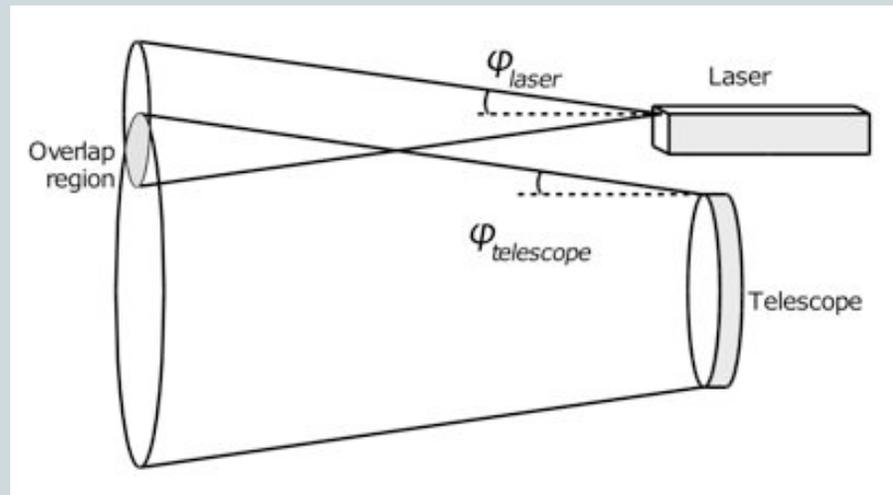
Modified Near-Field Detector Installation at Loma Amarilla



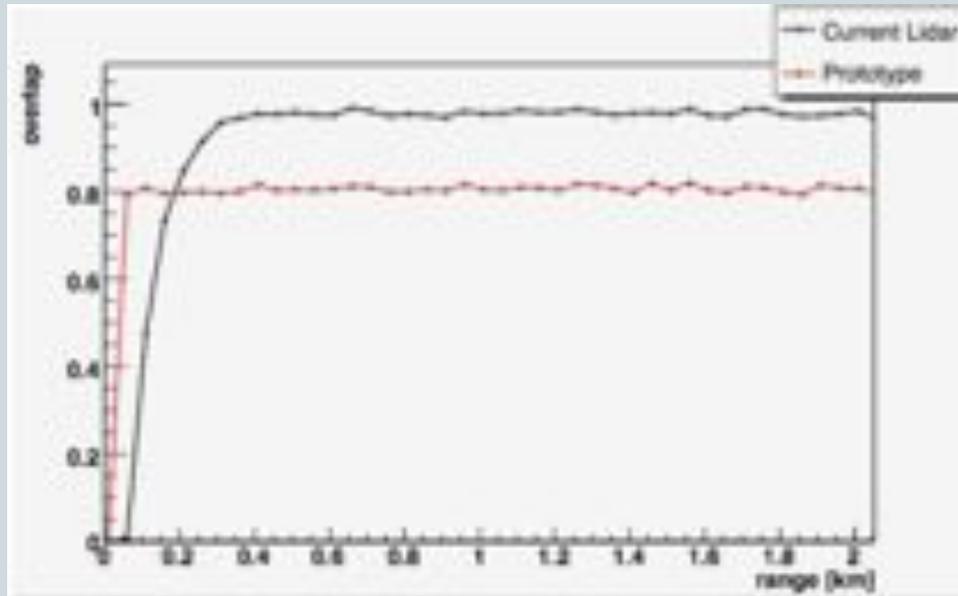
LIDAR Data: Overlap Function

- Data from the near-field LIDAR detector is analyzed to gather specific information about the atmosphere.
- Evaluate the Overlap function, which serves to indicate proper alignment of the telescope with respect to the laser.
- The overlap function $G(r)$ is a ratio of the effective receiving area of the detector to the total cross-sectional area of the back-scattered laser signal.

$$G(r) = A_{eff}(r) / A_0$$

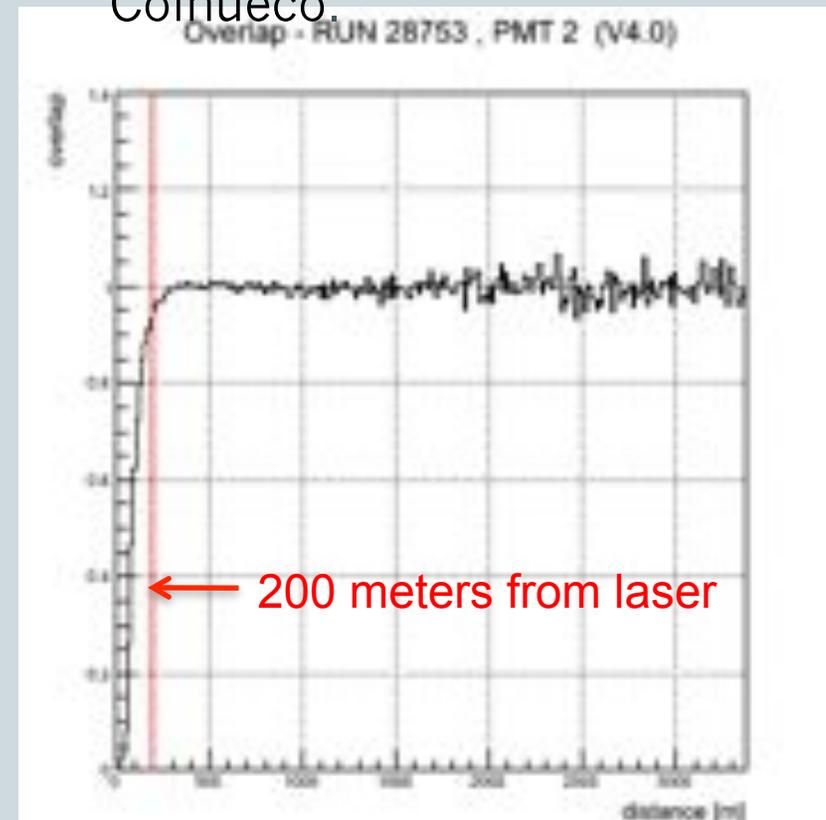


Performance Evaluation



- Comparison of simulated overlap functions for the original LIDAR detectors (black) and the near-field LIDAR detectors (red).
- Near-field detectors achieve complete overlap at much closer range.

- Overlap function generated from atmospheric data collected by the prototype near-field detector at Coihueco.

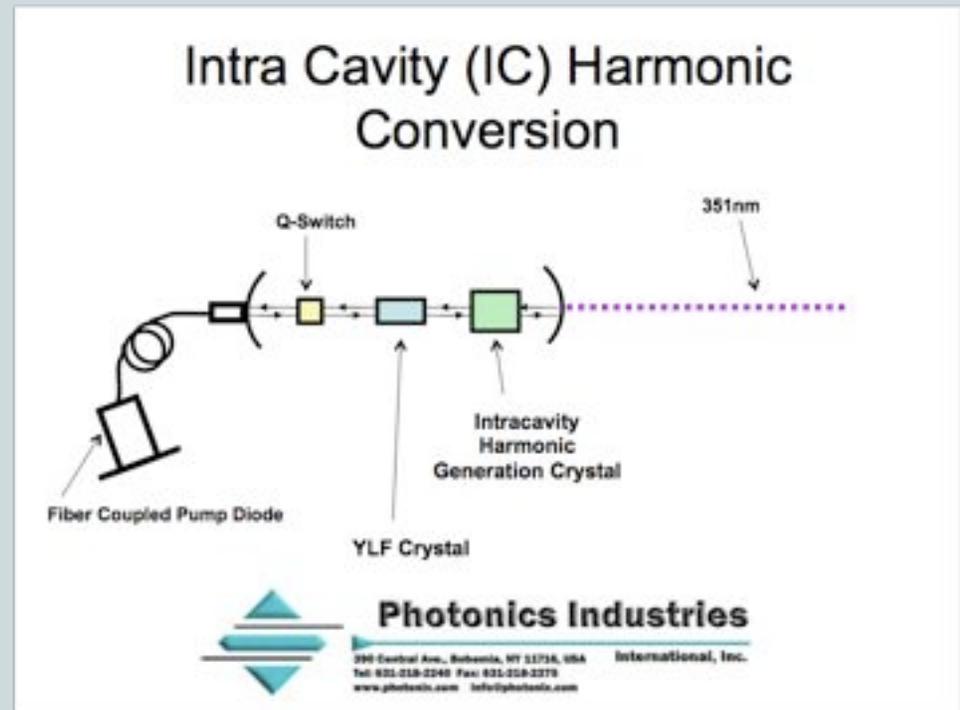


Future Work on Near-Field Upgrade

- Build and install two additional near-field detectors for the remaining fluorescence detector sites, Los Leones and Los Morados
- Analysis of data collected by near-field detectors
- Demonstrate that the addition of near-field atmospheric measurements improves the accuracy of cosmic ray energy calculations

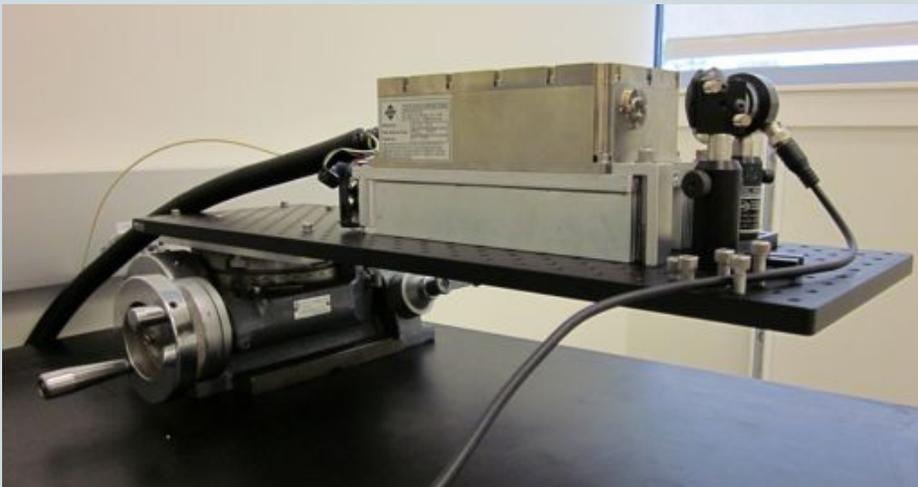
Laser Study

- Objectives:
 - Check stability of laser power over extended periods of operation
 - Check stability of laser power at varying physical orientations
- LIDAR Laser:
 - Diode-pumped, solid state laser: Photonics Industries model #DC30-351
 - 351 nm wavelength
 - Type Nd:YLF (gain medium is a neodymium-doped yttrium lithium fluoride crystal)
 - Average power: 300 mW
 - Average pulse energy: 150 μ J



Experimental Setup & Procedures

- Laser is operated at a rep. rate of 333 Hz to emulate actual operating conditions
- Use a mirror to align laser beam with the photodiode detection area
- Laser, mirror, and photodiode mounted on rotary table
- Photocurrent from photodiode is transferred to the oscilloscope via BNC cable
- Oscilloscope signal is the voltage measured across a 1 M Ω resistor
- Visual Basic program records laser pulse data for analysis
- 1-hour runs at each elevation angle of interest: 0, 30, 60, and 90 degrees
- 10-hour runs in the horizontal position



NASA Education Materials offer New Horizons for Online Physics Courses at UNL¹

Maria Becker², Dr. Gregory Snow³, and Suhasini Kotcherlakota⁴

Light, Color, and Their Uses



The Image Formed by a Concave Mirror

A concave mirror that is part of a ball or hollow sphere (that is, it has a circular cross section) is a spherical mirror. The focal length is approximately one-half the radius of curvature. A ray that is both parallel and very close to the optical axis will be reflected by the mirror so that it will cross the optical axis at the "paraxial focal point." The paraxial focal point is located a distance of one-half the radius of curvature from the point on the mirror where the optical axis intersects the mirror. The word "paraxial" comes from the Greek "para" or "near" meaning "in the side of, or beside, and axial." Thus paraxial means beside the axis.

Another ray that is parallel to the optical axis, but not close to the axis, will be reflected by the mirror so that it crosses the optical axis, not at the paraxial focus, but a small distance

closer to the mirror. This difference in the way cross-over points is called spherical aberration.

If the mirror has a cross section that is a parabola instead of a circle, all of the rays that are parallel to the optical axis will cross at the same point. Thus, a paraboloidal mirror does not produce spherical aberration. This is why the astronomical telescope, known as the Newtonian (named by Isaac Newton) uses a paraboloidal primary mirror.

For demonstrative purposes in the classroom, it works out that we can make the approximation that spherical mirrors behave almost like paraboloidal mirrors and determine that the focal length of a spherical mirror is about one-half the radius of curvature of the mirror.



Image: An Educator's Guide With Activities to Science and Mathematics, 11-122-11-14 (MS-1)

http://www.nasa.gov/audience/educators/epscor/materials/activities/type/Optics_Guide.html

NASA Educator Guides provide straightforward explanations and diagrams that demonstrate how mirrors and lenses bend and focus light. Optics is one of the key learning areas of the online course PHYS 212.



Department of Physics & Astronomy,
University of Nebraska- Lincoln

Introductory, Calculus-Based Physics Courses:

PHYS 211: Mechanics, Gravitation, Materials, and Waves

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Here we teach pilots and engineers how to evaluate aircraft—you know, see if they perform

http://www.nasa.gov/mov/100440main_016_intro_newton_laws.mov



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Video Learning Clips provide lecture-style education, not otherwise available with online courses.

These still shots from featured clips show a NASA test pilot explain Newton's Three Laws of Motion and an astronaut in orbit around Earth explain the force of gravity and how it applies to a tethered satellite in space.

Mechanics and gravity are central components of the PHYS 211 curriculum.



¹ Work supported by a NASA Nebraska Higher Education Mini-Grant. ² Maria Becker is an undergraduate Physics major at UNL. ³ Dr. Gregory Snow is an Associate Dean of the College of Arts & Sciences at UNL, as well as a Professor for the Department of Physics & Astronomy. ⁴ Suhasini Kotcherlakota is an Instructional Design Specialist with E&SO Instructional Design & Development at UNL.

Summary

- **Two NASA Nebraska Research Mini-Grants** has enabled the University of Nebraska to enhance the atmospheric monitoring capabilities of the Pierre Auger Observatory in Argentina
- The near-field LIDAR detectors (prototype, then improved design) have been shown to work as designed
- This project is half finished and will continue through 2010-2011
- A **NASA Nebraska Higher Education Mini-Grant** has enabled us to add new dimensions to two online Introductory Physics courses offered by UNL
- Comment on collaborations with NASA scientists (CALYPSO, NASA Langford Laboratory)