

Western Regional Space Grant Directors Meeting
Bend Oregon, September 2016

MODERN EDDINGTON EXPERIMENT (MEE)

Toby Dittrich

Portland Community College

Jack Higginbotham

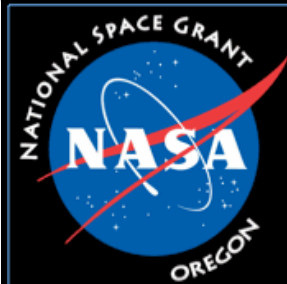
Oregon State University

Ralph Emerson

PlaneWave Inc



**Portland
Community
College**



Early
Academic
Interests



Interest in
Eclipses?

La Paz Mexico Eclipse 1991



[Youtube 1991 La Paz Eclipse](#)





Bellatrix

Tejat Posterior

Mebstata

Betelgeuse

Mintaka
Alnilam
Alnitak

Alhena

Castor

Pollux

Sun

Hats

S

Gomeisa

Procyon

Sirius

Mercury

Algenubi

Jupiter

o2

Hydrobius

ieba

Mars

Regulus
Venus

p Pup



Pollux

Mercury

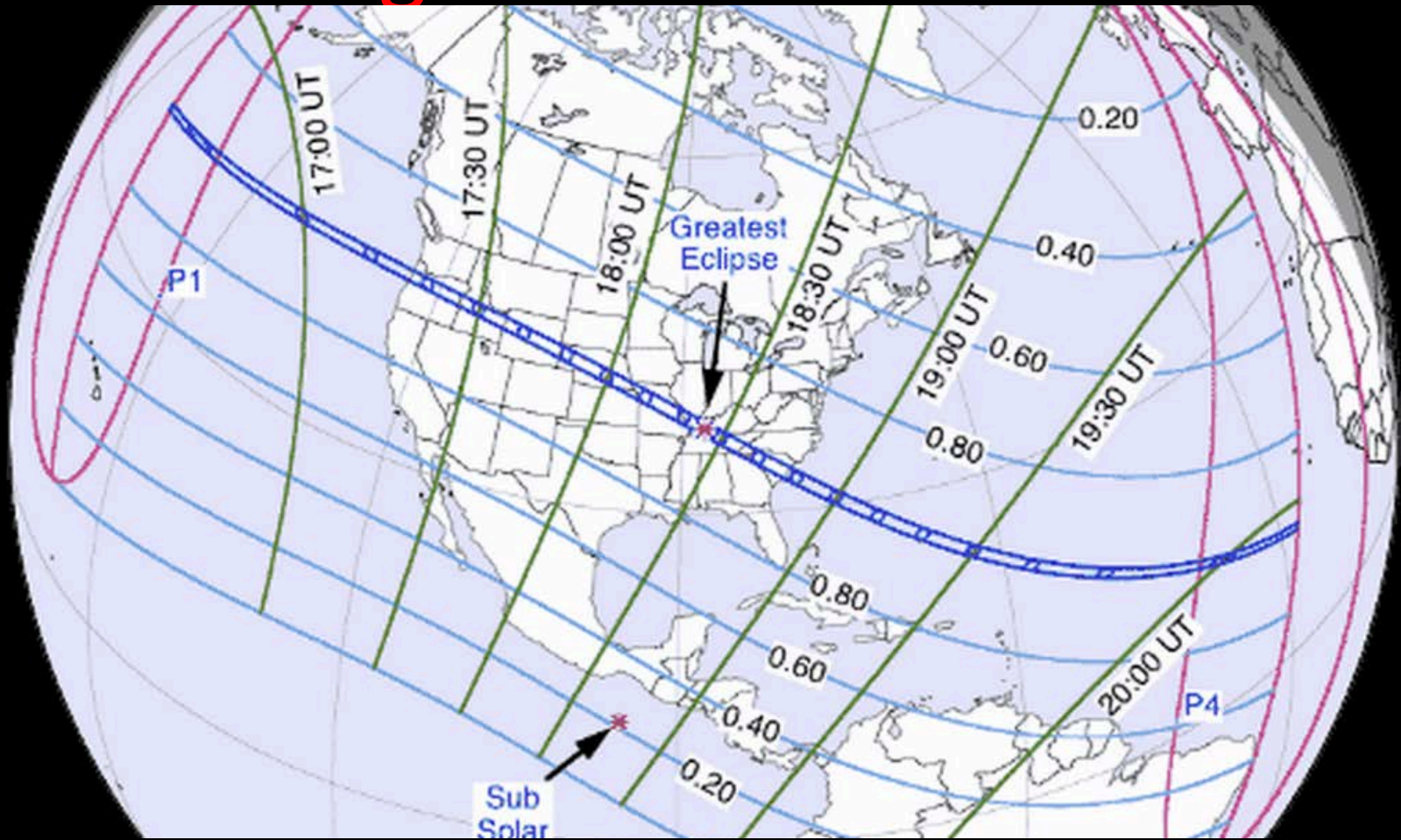
Jupiter

Mars

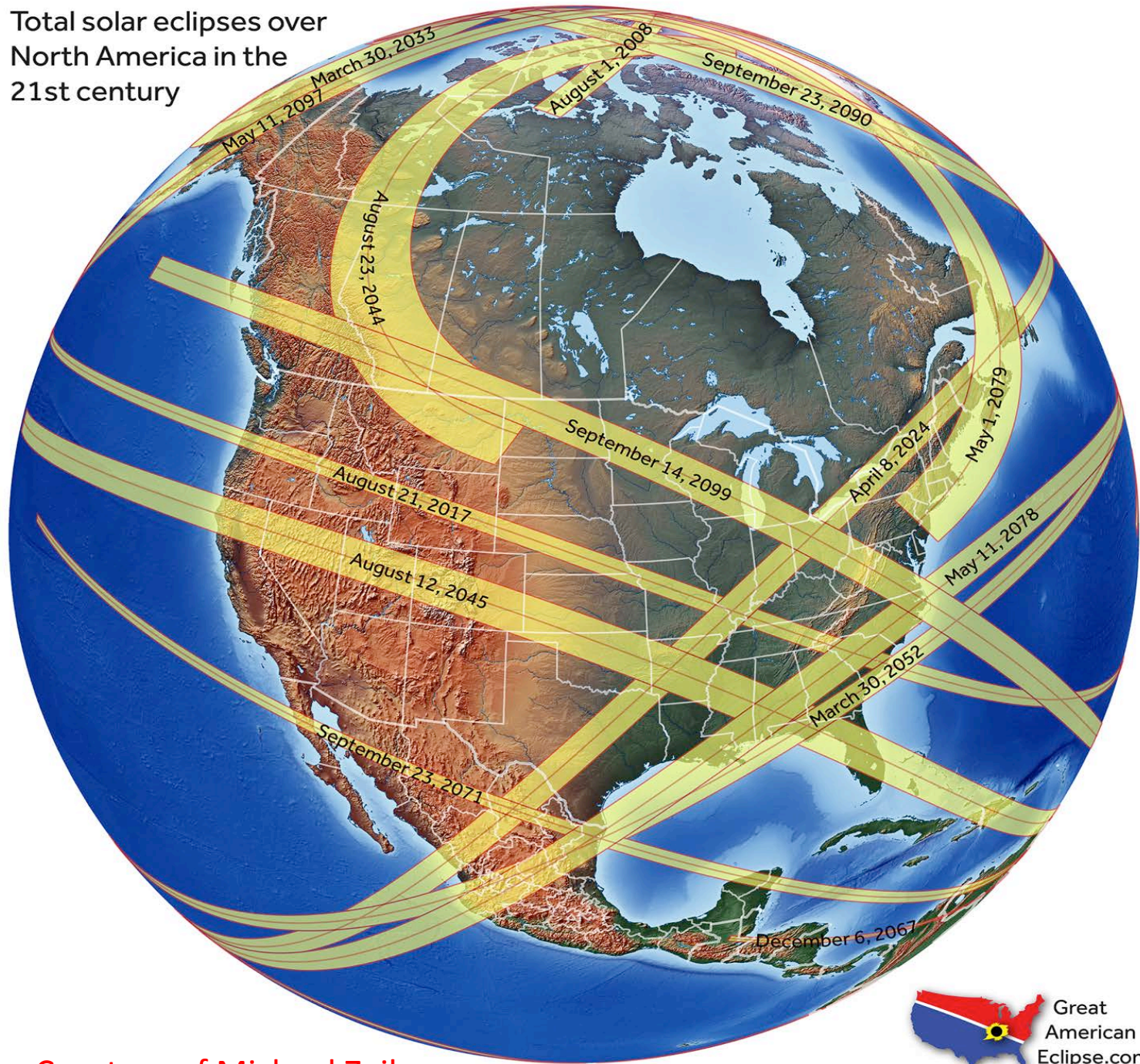
Regulus

Venus

August 21 2017



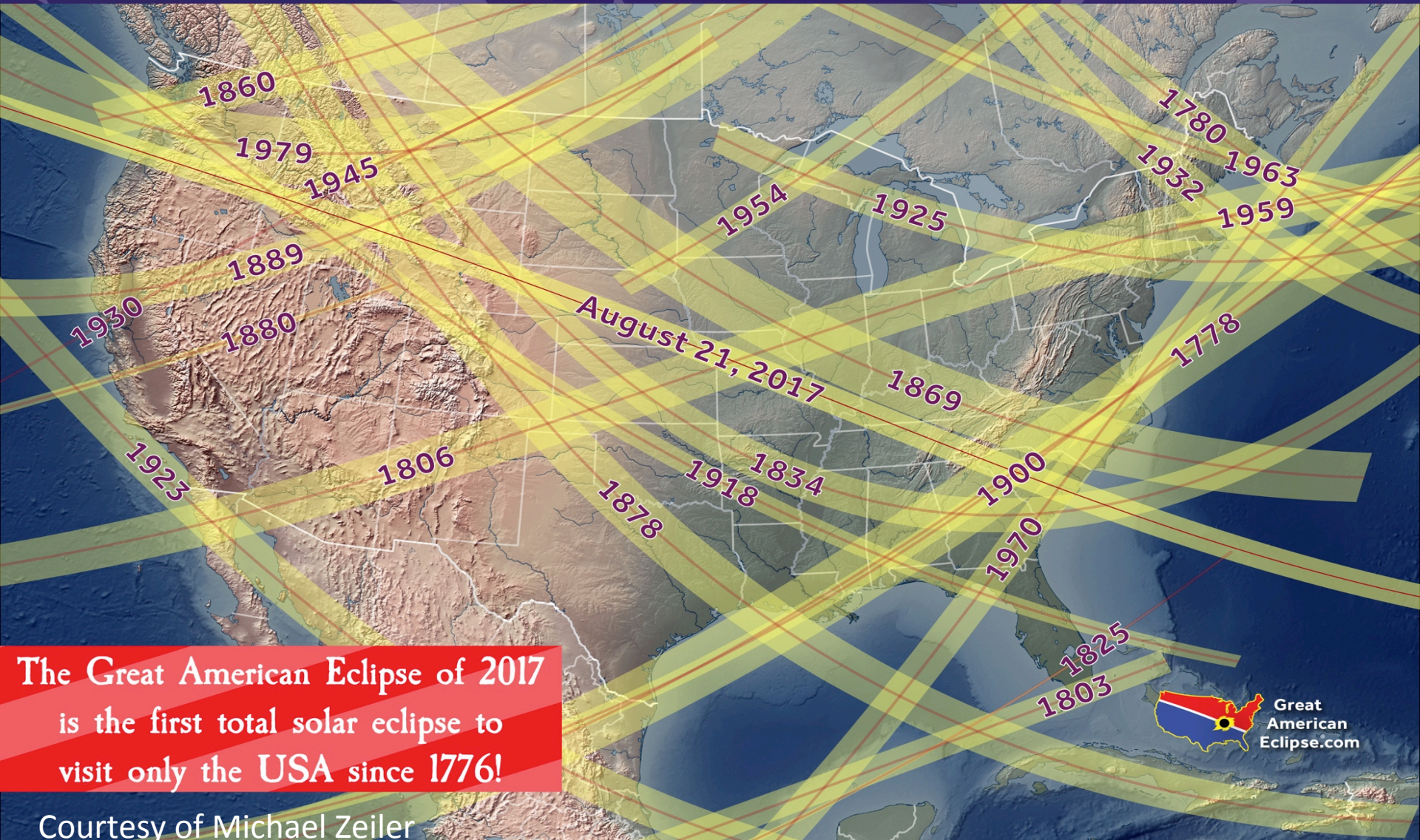
Total solar eclipses over North America in the 21st century



Courtesy of Michael Zeiler



All the total solar eclipses across the United States of America from 1776 to 2017



The Great American Eclipse of 2017 is the first total solar eclipse to visit only the USA since 1776!



Courtesy of Michael Zeiler

Heart of Constellation LEO



5 UMa

Praecipua
46 UMa 37 LMi

Rasalas Alterf
Algenubi

Alula Borealis
Alula Australis

Adhafera

Mars

54 Leo

Algieba
40 Leo

Al'dzhabkhakh

ξ Leo

72 Leo

b Leo

Sun
Regulus

Subra

Zosma

10 Leo

31 Leo
π Leo

2 Sex

93 Leo

Chertan

ρ Leo

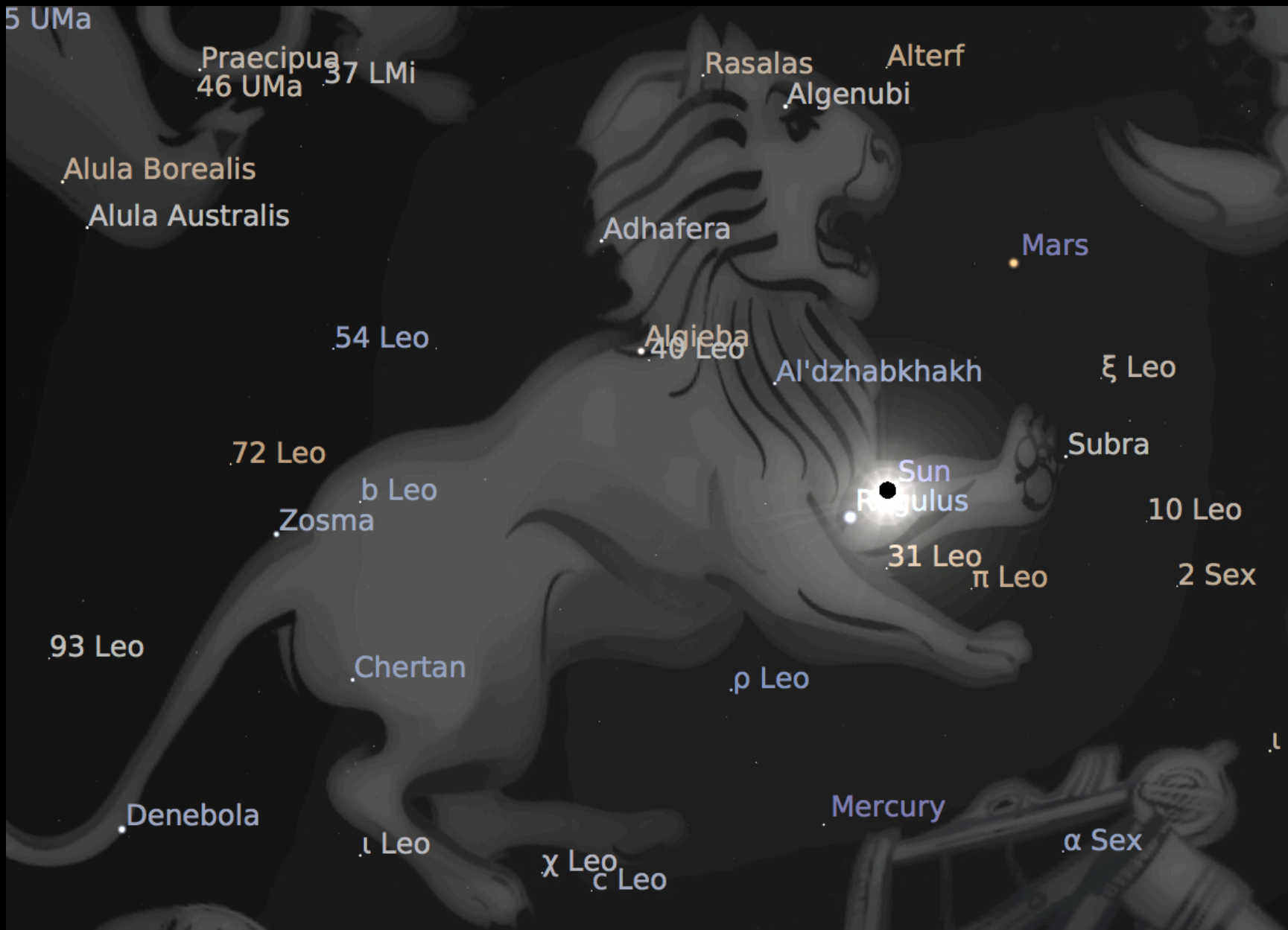
Denebola

ι Leo

χ Leo
c Leo

Mercury

α Sex



Animation of the shadow across USA

The Great American Eclipse
of August 21, 2017



9:56 a.m. PDT

10:56 a.m. MDT

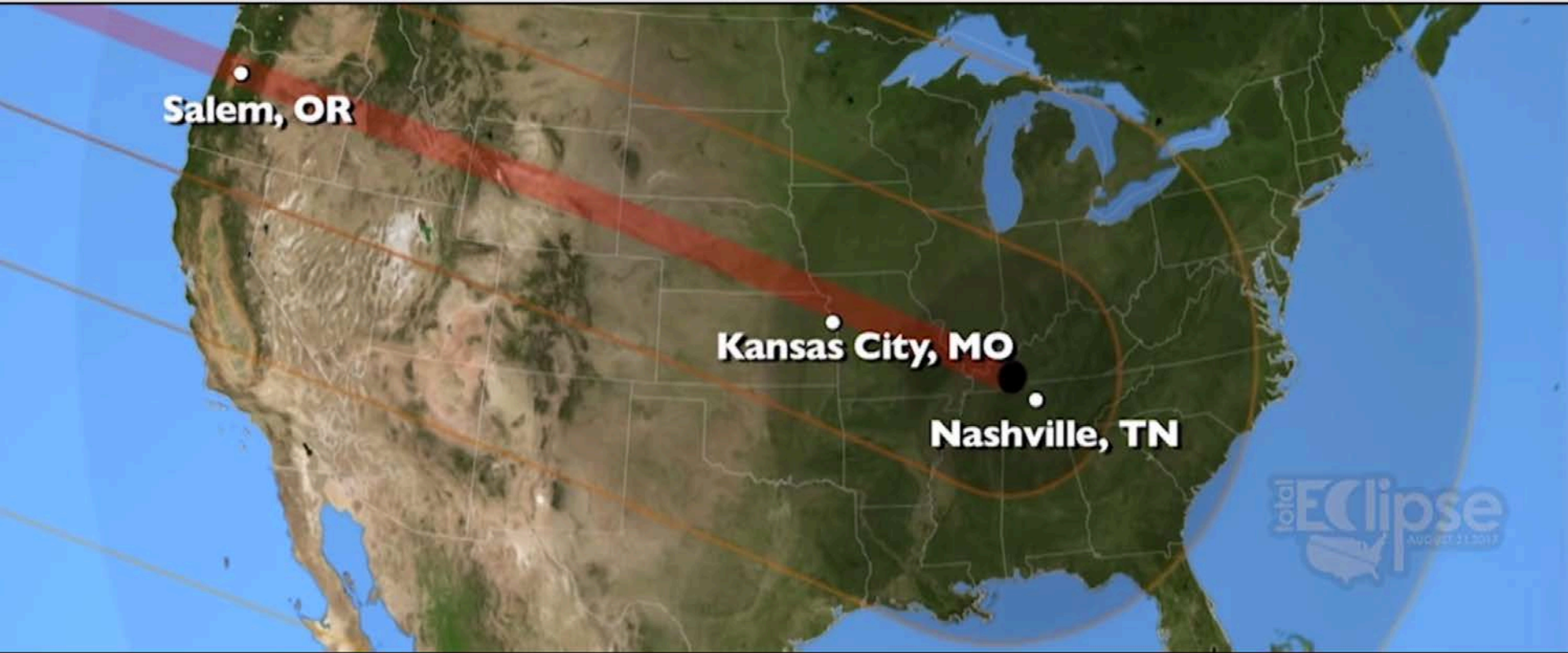
11:56 a.m. CDT

12:56 p.m. EDT

Courtesy of Michael Zeiler

National Space Grant Network

- States In path of Totality (or near)
- Oregon, Idaho, Wyoming, Nebraska, Iowa, Kansas, Missouri, Illinois, Kentucky, Tennessee, Georgia, North Carolina, South Carolina
- <http://eclipse2017.nasa.gov>



Eclipse Countdown Until First Contact in Oregon August 21, 2017 UT

47:00:14:05:52



<http://eclipse2017.nasa.gov/>



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Workspace design and creation by [Zing Studios](#). Graphic Design by [Ideum](#).

<http://smdepo.org/group/6162>

 Reed, Shannon P. (GSFC-670.0)[ADNET SYSTEMS INC]

Sep 14 (11 days ago) ☆

to nasa-eclipse20. ▾

Good Morning,

Just a reminder that we will be holding our monthly (3rd Friday of the month) **Eclipse 2017 Telecon** Friday, September 16, 2016 at 3:00pm - 4:00pm (EDT).

AGENDA

-Eclipse Website Update

<http://eclipse2017.nasa.gov>

-Call for Challenges, Citizen Science and DIY Science Projects

-Eclipse Glasses and Eclipse in a Box update

-Safe Solar Viewing message

-SMDEPO Reminder

<http://smdepo.org/group/6162>

-Open Floor

Eclipse Resource Guide for
teachers, students, and the public at
the nonprofit Astronomical Society
of the Pacific website:
<http://www.astrosociety.org/eclipse>

Telecon call-in information:

USA Toll Free #: [1-844-467-4685](tel:1-844-467-4685)

USA Local/Toll #: [1-720-259-7012](tel:1-720-259-7012)

Participant Passcode: #640276

Research Team

- Daniel Kennifick, University of Arkansas
- Don Bruns, StellarProducts.com
- Brad Schaefer, Louisiana State University

RESEARCH TEAM



feature
article

Testing relativity from the 1919 eclipse— a question of bias

Daniel Kennefick

When interpreting experimental results, context is everything. The researchers who took and analyzed the most important eclipse data had good reasons for judging the experiment a victory for Albert Einstein.

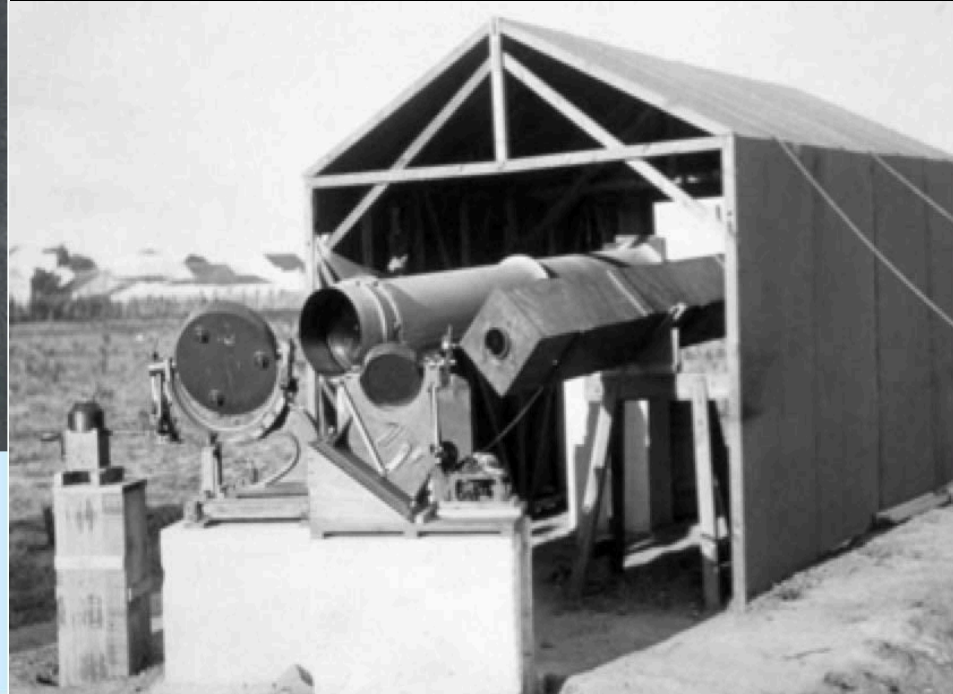
Daniel Kennefick is an assistant professor of physics at the University of Arkansas at Fayetteville.

RESEARCH

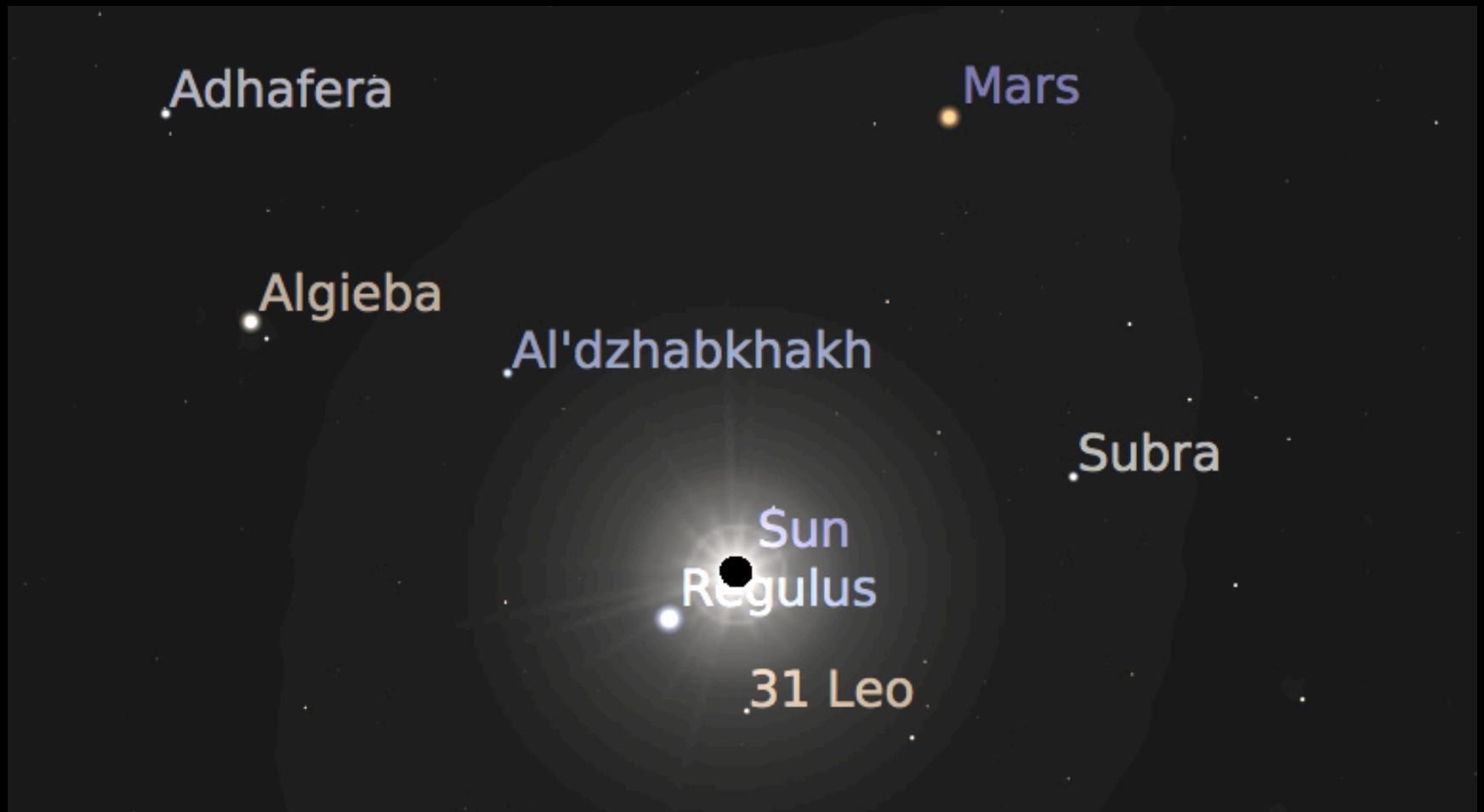
- 1st Attempt – Lindeman & Lindeman
August 21, 1917 India
- 2nd Attempt – Eddington & Dyson, Crimea
- 3rd Attempt – Eddington & Dyson 1919
Africa and Brazil



Arthur Stanley Eddington (1882–1944). In 1919 Eddington had already acquired a major reputation as a result of his work on the internal structure of stars. His enthusiasm for general relativity has led some historians to accuse him of bias in the analysis of the 1919 eclipse data. (Courtesy of the AIP Emilio Segrè Visual Archives.)



100th Anniversary August 21 2017



23 Leo

HIP 49445

.18 Leo

.19 Leo

Date and Time ✕

Date and Time Julian Date

2017 / 8 / 21 10 : 23 : 38

HIP 48298

HIP 49623

HIP 48033

DY Leo

HIP 49348

y Leo

HIP 49158



HIP 49518

HIP 48911

34 Leo

GM Leo

HIP 48921

Regulus

HIP 49925

HIP 48734

HIP 49279

HIP 50473

Eclipse Close Field View

Date and Time		Julian Date	
2017	8 / 21	10	5 : 35

DY



Star Positions

Mag	Star	RA	Dec	$d\theta$	r	Deflection
	Sun	10h 4m 3.9s	+11d 51' 43"			-----
1.35	Regulus	10h 9m 16.8s	+11d 52' 54"	1.2762°	4.79	0.365"
4.35	31 Leonis	+171 47'	+60 33'			
5.25	nu Leonis	+176 31'	+63 12'			
7.1	HIP 49328	10h 5m 32s	+11d 32' 39"	0.3995°	1.5	1.167"
7.75	HIP49158	10h 2m 51s	+12d 09' 41"	0.4216°	1.58	1.106"

Don Bruns



Don Bruns

Presented at the Society for Astronomical Sciences Symposium, June 17, 2016 in Ontario, CA

Measuring Starlight Deflection during the 2017 Eclipse: Repeating the Experiment that made Einstein Famous

Donald Bruns
7387 Celata Lane
San Diego, CA 92129
dbruns@stellarproducts.com

Abstract

In 1919, astronomers performed an experiment during a solar eclipse, attempting to measure the deflection of stars near the sun, in order to verify Einstein's theory of general relativity. The experiment was very difficult and the results were marginal, but the success made Albert Einstein famous around the world. Astronomers last repeated the experiment in 1973, achieving an error of 11%. In 2017, using amateur equipment and modern technology, I plan to repeat the experiment and achieve a 1% error. The best available star catalog will be used for star positions. Corrections for optical distortion and atmospheric refraction are better than 0.01 arcsec. During totality, I expect 7 or 8 measurable stars down to magnitude 9.5, based on analysis of previous eclipse measurements taken by amateurs. Reference images, taken near the sun during totality, will be used for precise calibration. Preliminary test runs performed during twilight in April 2016 and April 2017 can accurately simulate the sky conditions during totality, providing an accurate estimate of the final uncertainty.

Telescope: Tele Vue NP101

Tele Vue-NP101is



TeleVue.com: Telescopes > Tele Vue-NP101is > Home

101mm, f/5.4 APO (Nagler-Petzval) Refractor - Imaging/Visual.

Based on exemplary wide field and planetary visual scopes with added features to make imaging easier and more flexible.

4.9° / 250x (Max. Visual Field / Max. Rec. Power)



TV-NP101is image
Heart & Soul Nebula with
STL11000M camera
(Fabian Neyer)



- 101 mm aperture
- capture 10th magnitude stars with 1 second exposures
- diffraction limit is only 1.3 arcsec at 630 nm

*“Greater than expected! Pleasure to use,
excellent optics. I don't need XXX no more.”*

—S.Y., Japan

Camera: Microline 8051 CCD

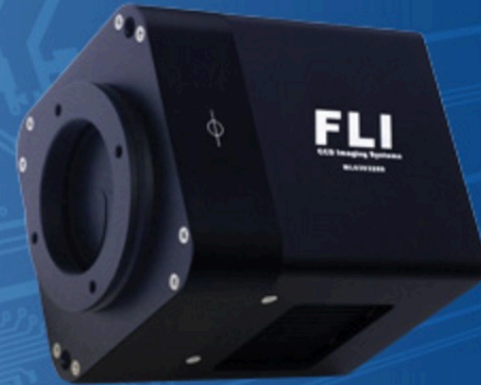
MicroLine

Low Noise

High Sensitivity -8 MPixel 8051model

Deep Cooling -12MHz digitizing speed

Small Footprint



MicroLine Cameras

Goodbye Goliath! MicroLine is smaller and lighter-weight than the competition, but outperforms them in key areas such as noise, frame quality, download speeds, bit-depth, and CCD cooling.

- Sensor Cooling up to 65°C below ambient (air cooled)
- Interline sensor operation at 12 MHz (16-bit)
- Full frame sensor operation up to 12 MHz (16-bit)
- Dual Super Coated AR Window - VIEW REFLECTANCE INFORMATION
- Heavy weight performance in a 2.8 lbs package
- Liquid or Air Cooled base
- RBI Anti-Ghosting Technology standard
- Video mode with interline sensors - focus and center images live!
- Small Footprint (3.7 x 3.7 x 5in.)
- Reaches operating temperatures in 5 minutes!
- Single, Dual, or Quad Output (4 Ch) Capable**

-pixels are only 5.5
microns wide
-2.1" per pixel
-2° by 2° field

Mount



Figure 5. The Software Bisque MyT Paramount and matching tripod is portable, allowing a change in eclipse sites in case of bad weather. The periodic error, after correction, is sub-arcsecond, necessary for high quality images.

Anticipated Data

Short Exposure- 200 ms

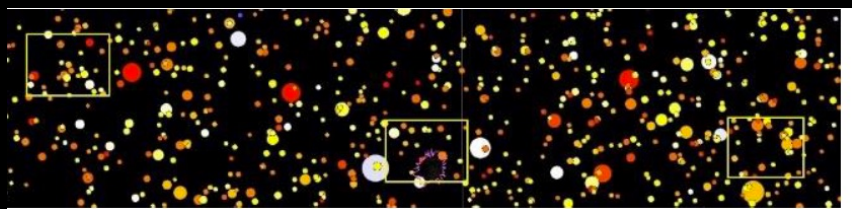


Long Exposure- 1 sec



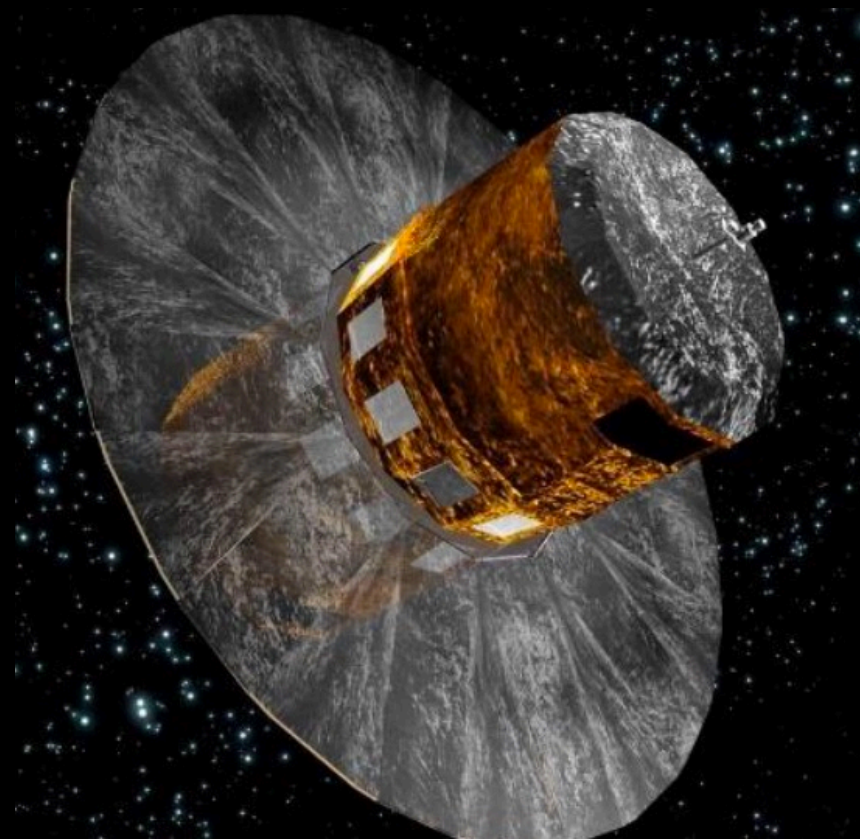
Additional Considerations

Calibration Reference Fields



0.000024 arcseconds.

Stellar Positions – Hipparchus Gaia



Bruns Expected Results

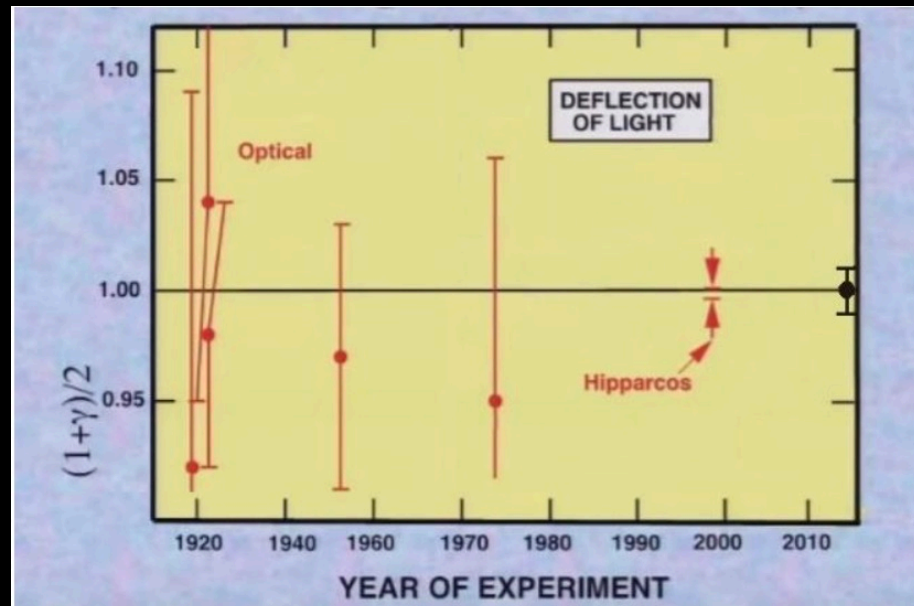


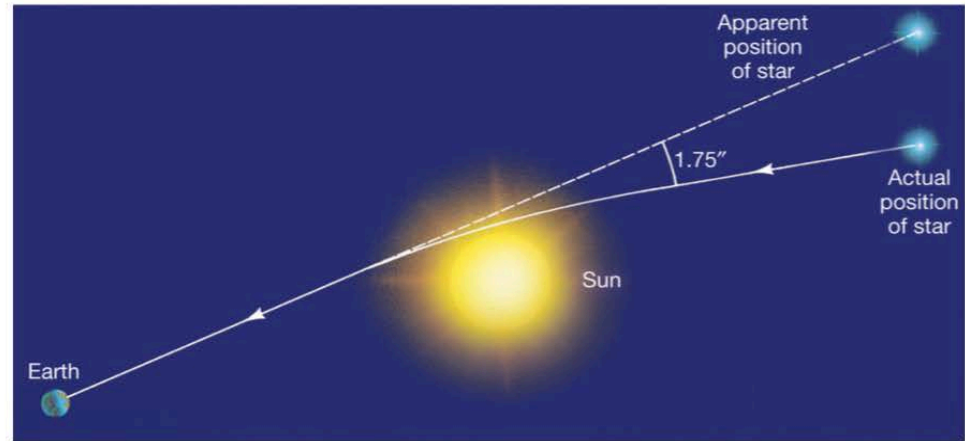
Figure 2. The deflection of light has been measured during a total eclipse only a few times in the last century. Hipparcos measured the deflection from space, not needing an eclipse, and has set the record for the most precise optical measurement. The marker on the right side of the graph represents the expected error of 1% for my 2017 experiment. Modified from [Will 2015]. Vertical scale is deflection normalized to Einstein's calculation.

Dr. Brad Schaefer, LSU

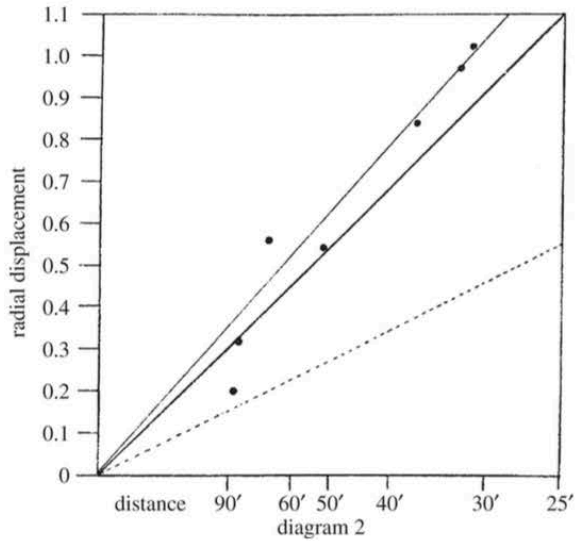


THE MODERN EDDINGTON EXPERIMENT

Brad Schaefer (BRAS)



EINSTEIN TRIUMPHANT



LIGHTS ALL ASKEW IN THE HEAVENS

Men of Science More or Less
Agog Over Results of Eclipse
Observations.

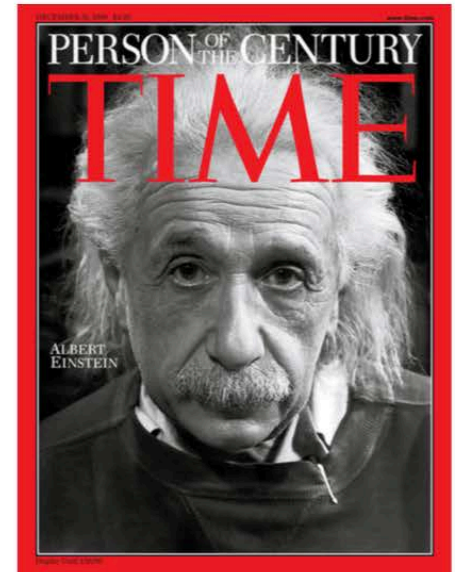
EINSTEIN THEORY TRIUMPHS

Stars Not Where They Seemed
or Were Calculated to be,
but Nobody Need Worry.

A BOOK FOR 12 WISE MEN

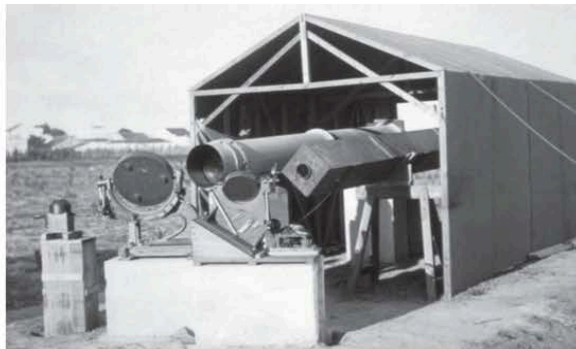
No More in All the World Could
Comprehend It, Said Einstein When
His Daring Publishers Accepted It.

Special Cable to THE NEW YORK TIMES.
LONDON, Nov. 9.—Efforts made to
put in words intelligible to the non-
scientific public the Einstein theory of
light proved by the eclipse expedition
so far have not been very successful. The
new theory was discussed at a recent
meeting of the Royal Society and Royal
Astronomical Society. Sir Joseph Thom-
son, President of the Royal Society, de-
clares it is not possible to put Einstein's
theory into really intelligible words, yet
at the same time Thomson adds:
"The results of the eclipse expedition
demonstrating that the rays of light
from the stars are bent or deflected
from their normal course by other aerial
bodies acting upon them and conse-
quently the inference that light has
weight form a most important con-
tribution to the laws of gravity given us
since Newton laid down his principles."



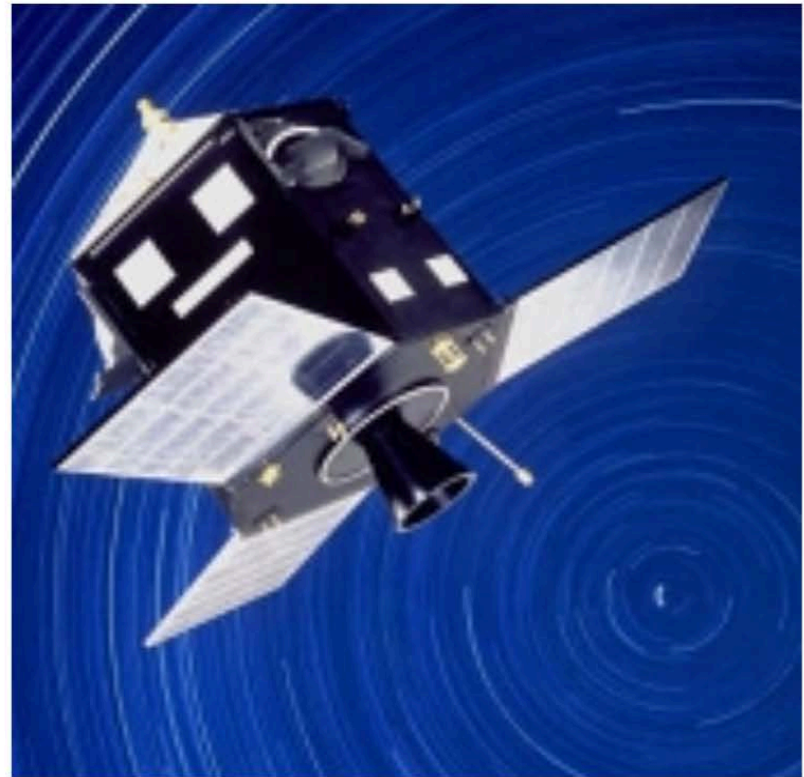
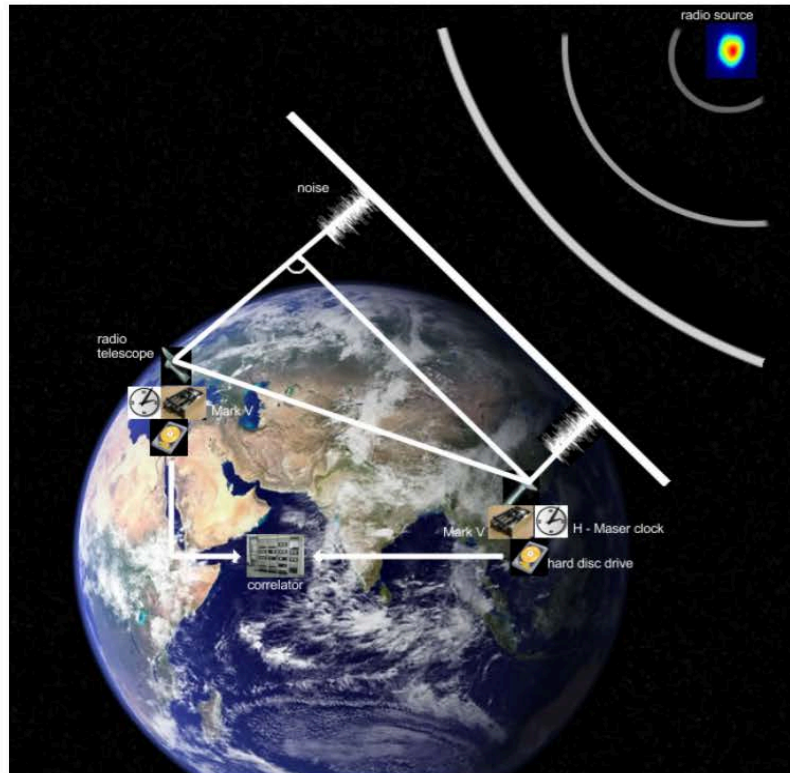
EDDINGTON EXPERIMENT AT OTHER SOLAR ECLIPSES

<u>ECLIPSE</u>	<u>RESULT</u>
1914 Crimea	Failed due to outbreak of WWI
1918 Goldendale Wash	Confused, maybe 0.87"
1922 Australia	1.72" \pm 0.11", 1.74" \pm 0.3", 1.77" \pm 0.3"
1926 Sumatra	Failed
1929 Sumatra	1.75" \pm 0.13"
1936 Russia, Japan	1.70" \pm 0.40", 2.73" \pm 0.31"
1947 Brazil	2.01" \pm 0.27"
1952 Khartoum	1.70" \pm 0.10"
1973 Mauritania	1.66" \pm 0.18"



LIGHT BENDING BY OTHER METHODS

1995	VLBI (3C 273 & 3C 279)	0.34%
1997	<i>Hipparcos</i> (many optical stars far from Sun)	0.20%
2004	VLBI (541 radio sources)	0.045%
2009	VLBI (4 radio sources)	0.030%
2010	VLBI (541 radio sources)	0.012%



CRITICAL ADVANCE IN THE LAST TWO YEARS

★ The sudden availability of commercial CCDs with 4096 pixels across



Note: we cannot use color CCDs, CMOS detectors, or Bayer pattern RGB pixelation.

BASELINE SYSTEM & PROCEDURE:

- ★ 6-inch f/5 scope (F=800 mm)
- ★ Polar-aligned mount with good tracking and 'goto' ability
- ★ SBIG STX-16803 (4096x4096, 9μ pixels, 100,000 e- full well)
- ★ Prior nights, take night sky images to practice and to map coma
- ★ During partial phase, focus scope on Pleiades, taken images
- ★ During totality, take ~30 two-second exposures
- ★ Immediately after totality, take Pleiades images

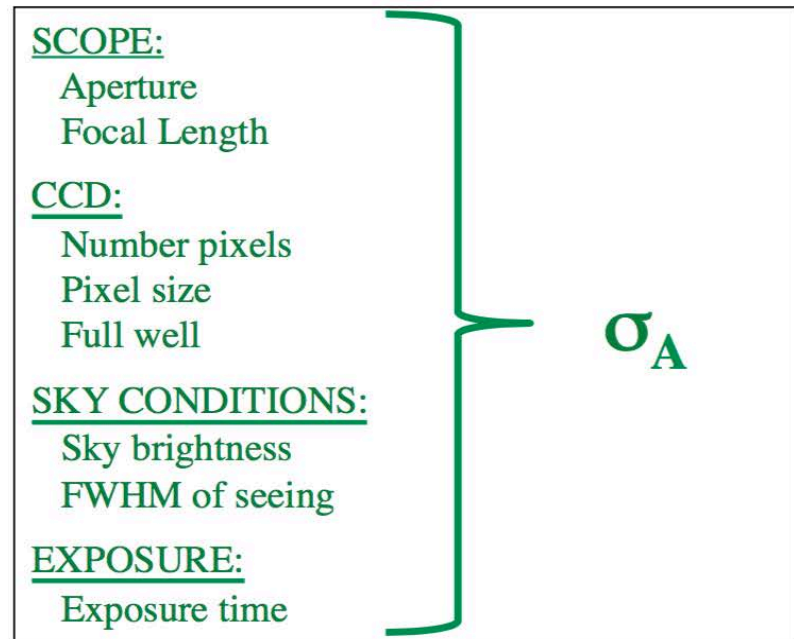


NEEDED: A FULL DETAILED CALCULATION OF EXPECTED ACCURACY IN MEASURING THE SHIFT

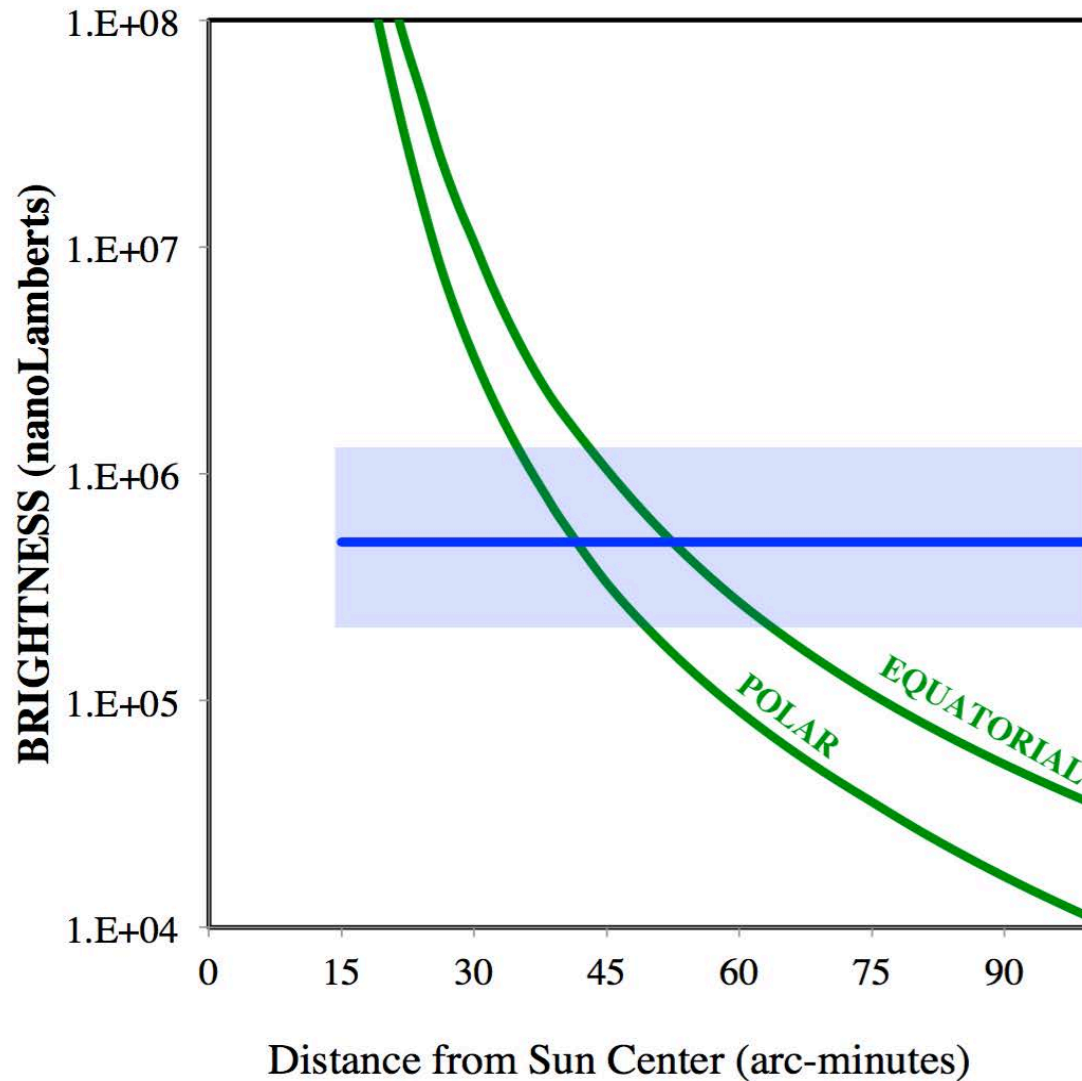
Model for the shift: $\Delta\Theta = A * 1.75'' (15'/\Theta)$

A=0.0 no shift
A=0.5 Newtonian case
A=1.0 GR prediction

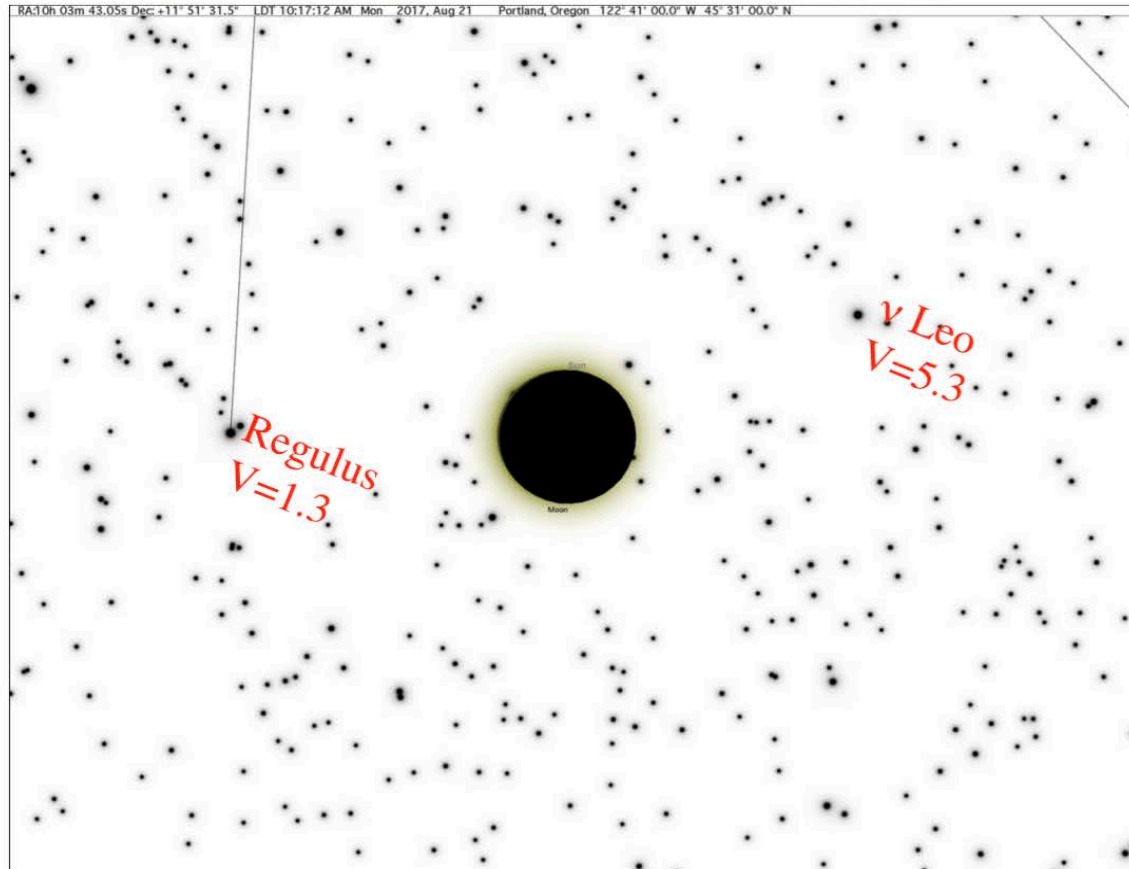
- ★ Corona brightness
- ★ Sky brightness
- ★ Star catalog
- ★ e- on CCD from stars
- ★ e- on CCD from background
- ★ Centroid accuracy
- ★ Calculate σ_A



SKY & CORONA BRIGHTNESS



STAR CATALOG



STAR	Θ (')	V (mag)	$\Delta\Theta_{GR}$ (")
Tyc 0832-0973-1	17.6	11.65	1.49
Tyc 0832-1033-1	21.5	10.99	1.22
Tyc 0833-0457-1	22.8	11.12	1.15
HD 86898	23.1	7.80	1.14
Tyc 0832-0276-1	23.6	11.20	1.11
Tyc 0833-0812-1	23.8	11.64	1.10
Tyc 0832-0272-1	24.9	11.74	1.05
HD 87271 = GM Leo	26.1	7.13	1.01

e- ON CCD

$$B_{\text{bck}} = B_{\text{corona}} + B_{\text{sky}}$$

$$\Theta_{\text{pixel}} = (L_{\text{pixel}}/1000) * 206264/F$$

$$\Theta_{\text{edge}} = 0.5 * N_{\text{pixel}} * (L_{\text{pixel}}/1000) * 206264/F$$

$$\text{Total Electrons On CCD in V-band} = E_{\text{star}} = (150,000 \text{ e-}) * D^2 * \Delta T * 10^{-0.4V}$$

$$\text{Total Electrons On CDD over 1 sq-arc-sec} = E_{\text{bck}} = (4.4 \times 10^{-6} \text{ e-}) * B_{\text{bck}} * D^2 * \Delta T$$

$$\text{Pixel Saturated if } (E_{\text{star}} + E_{\text{bck}}) * \Theta_{\text{pixel}}^2 > E_{\text{max}}$$

$$\text{Star detected if SNR} = E_{\text{star}} / (E_{\text{star}} + E_{\text{bck}} * \pi * \Theta_{\text{star}}^2)^{0.5} > 5$$

D = scope aperture (cm)
 F = focal length (mm)
 L_{pixel} = pixel size (μ)
 N_{pixel} = pixels on a side (#)
 E_{max} = full well (e-)
 ΔT = exposure time (s)
 Θ_{star} = FWHM (")
 B_{sky} = sky brightness (nL)
 B_{corona} = corona bright. (nL)
 V = star magnitude (mag)
 Θ = offset from Sun center (")

CENTROID ACCURACY

$$\sigma_{\text{centroid}} = [(FWHM/2.36/SNR)^2 + (\Theta_{\text{pixel}} f / SNR)^2]^{0.5}$$

$$f = 1.3 + 0.8 * [(FWHM/2.36/\Theta_{\text{pixel}}) - 1]^2$$

ACCURACY FOR EINSTEIN SHIFT

$$A_{\text{best}} = \Sigma [(\Delta\Theta_i * 1.75'' * (15'/\Theta)) / \sigma_{\text{centroid}}]^2 / \Sigma [1.75'' * (15'/\Theta) / \sigma_{\text{centroid}}]^2$$

$$\sigma_A = (A_{\text{best}} * \Sigma [1.75'' * (15'/\Theta) / \sigma_{\text{centroid}}]^2)^{-0.5}$$

SYSTEMATIC PROBLEMS & SOLUTIONS

★ Too few stars recorded.

CCDs go deep with adequate exposure

Require fields of view larger than $\sim 2^\circ$

★ PlateScale/GRshift degeneracy.

Image Pleiades (or Praesepe) immediately before/after totality

Have large field of view (larger than $\sim 2^\circ$)

★ Coma in scope can cause uncontrolled radial shifts.

Image Pleiades (or Praesepe) immediately before/after totality

Images of night sky during nights before and after

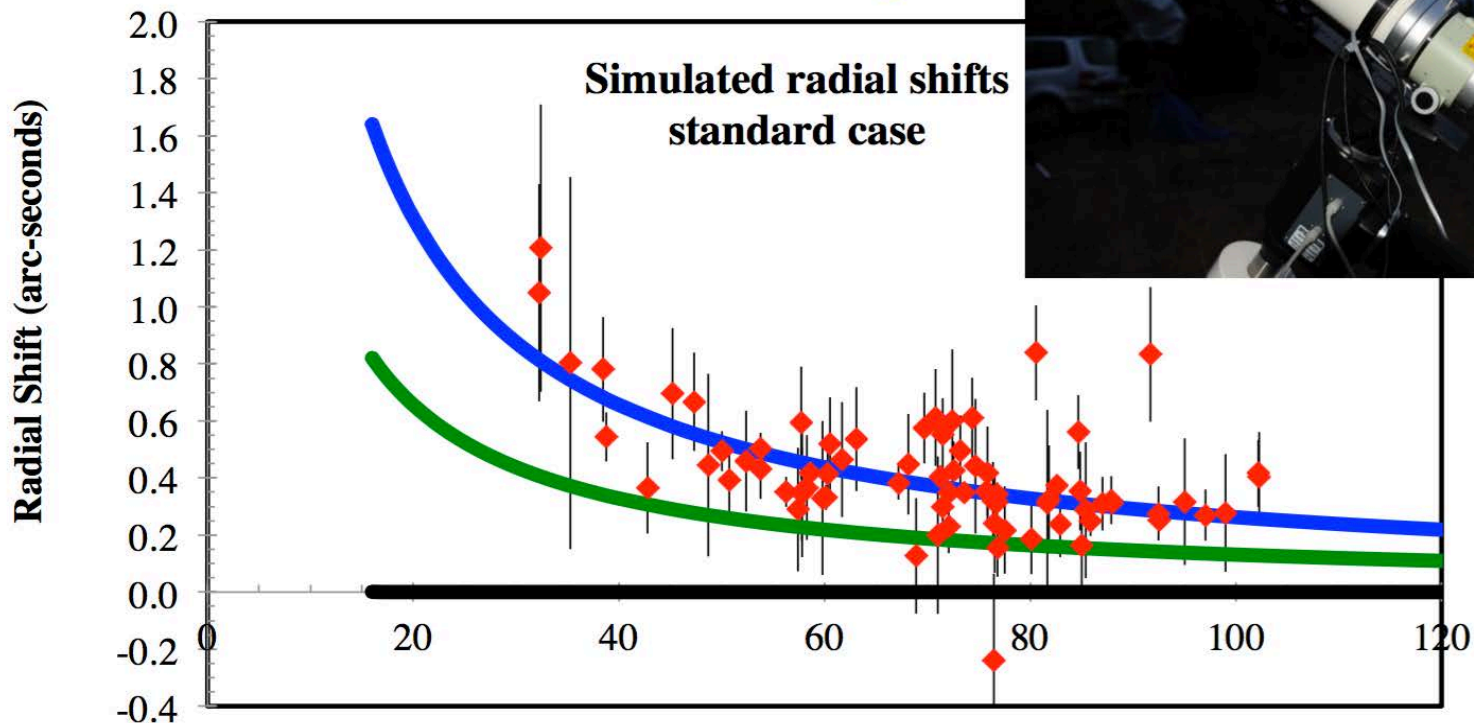
RESULTS

★ One person:

ONE IMAGE: 1.0% accuracy, better than all prior expeditions
FROM ALL ECLIPSE: 0.19% accuracy - WOW

★ Many people:

~100 PEOPLE: 0.019% accuracy



Summary

- Why perform the MEE? Win the ROSES SMD grant?
- Space Grant Involvement
- Preparation for 2024 eclipse
- Fabulous opportunity for Faculty, Students and Citizen Scientists
- tdittric@pcc.edu toby.dittrich@gmail.com

QUESTIONS?

The Great American Eclipse
of August 21, 2017



9:56 a.m. PDT

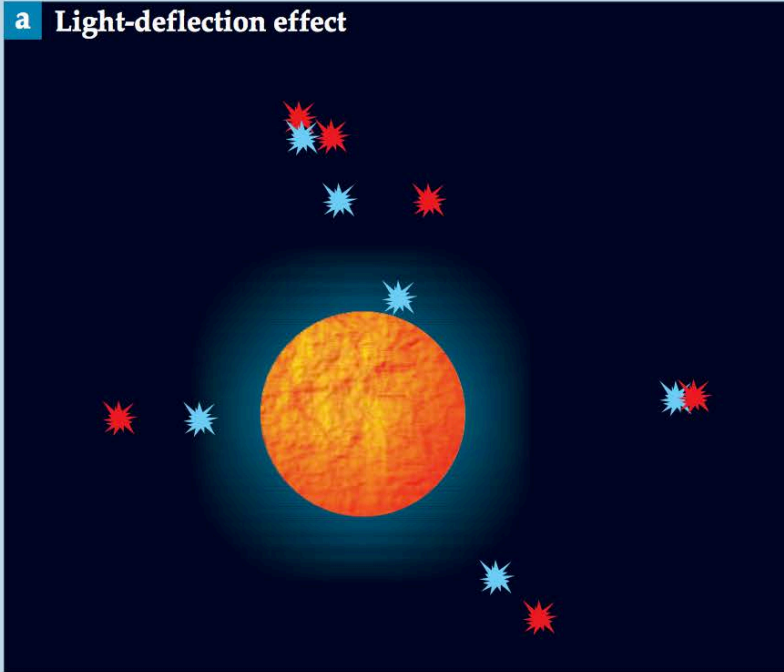
10:56 a.m. MDT

11:56 a.m. CDT

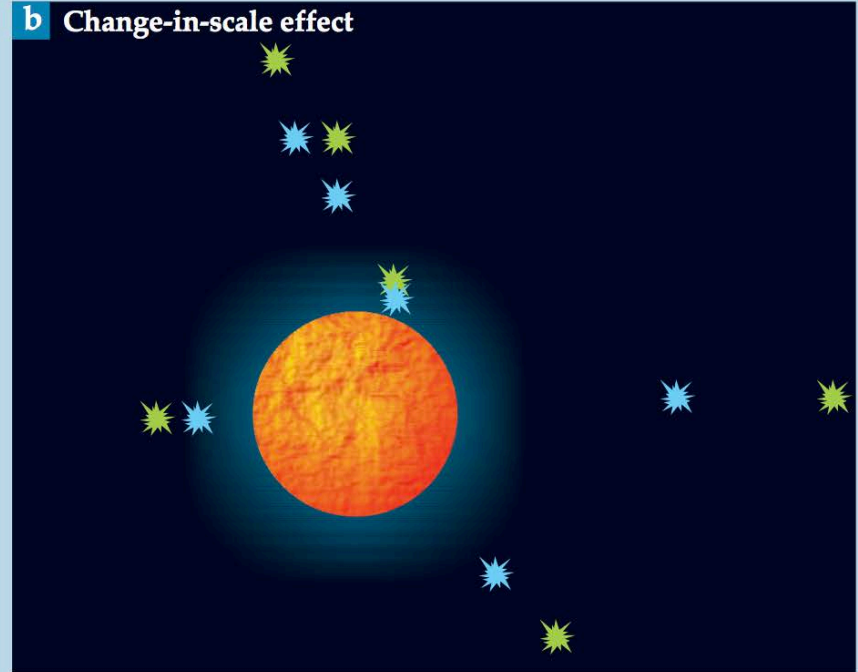
12:56 p.m. EDT

Temperature Magnification Effect

a Light-deflection effect



b Change-in-scale effect



An unusual number of bright stars from the Hyades cluster conveniently filled the sky on all sides of the Sun during the 1919 eclipse. **(a)** In this artist's rendering, deflection of starlight from the Sun's gravity shifts the original stellar positions (blue) radially away from the center of the Sun; those closest to the center suffer the greatest shift. **(b)** A change in scale between one exposure and another shifts stellar positions radially away from the center of the image; those far from the center show the greatest shifts.